

Topological Dependence of Runaway Avalanche Threshold in Momentum Space

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Outline

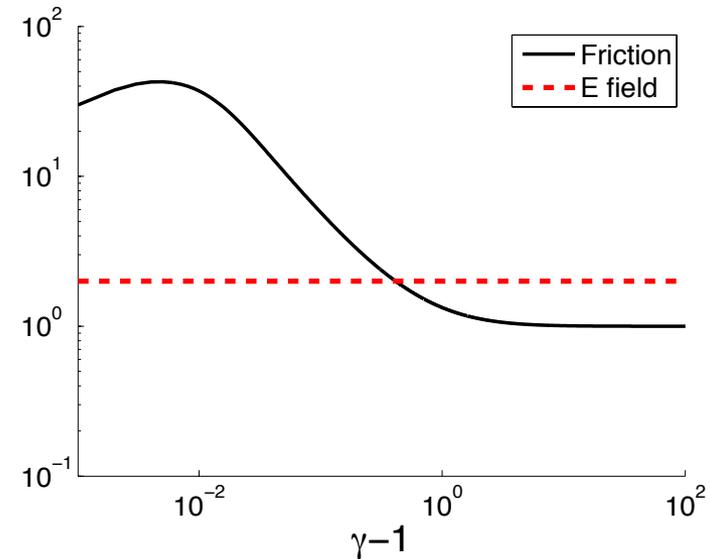
- Runaway Vortex: Momentum space topology of primary runaway distribution function
- Link of avalanche threshold to momentum space topological change \Leftrightarrow O-X merger (disappearance of runaway vortex)
- Self-consistent formulation of small and large angle collisions \rightarrow Impact on avalanche threshold and growth rate

Motivation: Avalanche Threshold

- The avalanche threshold is a crucial quantity for disruption mitigation scenarios
 - Provides guidance as to what conditions are necessary in order to avoid runaway generation through secondary generation
- An accurate prediction of the avalanche threshold has proven difficult:
 - Large amounts of high-Z materials typically present \Leftrightarrow resulting collisional coefficients highly complex (free-bound, scattering by partially shielded nucleus, etc) [[Breizman IAEA 2016](#), [Hesslow et al. 2017](#)]
 - Runaway models typically make different modeling assumptions \Leftrightarrow leads to order one variation of avalanche threshold [[Rosenbluth-Putvinski 1997–Martin-Solis 2017](#)]
- Here we are interested in identifying a robust indicator of the avalanche threshold
 - Independent of specific model assumptions

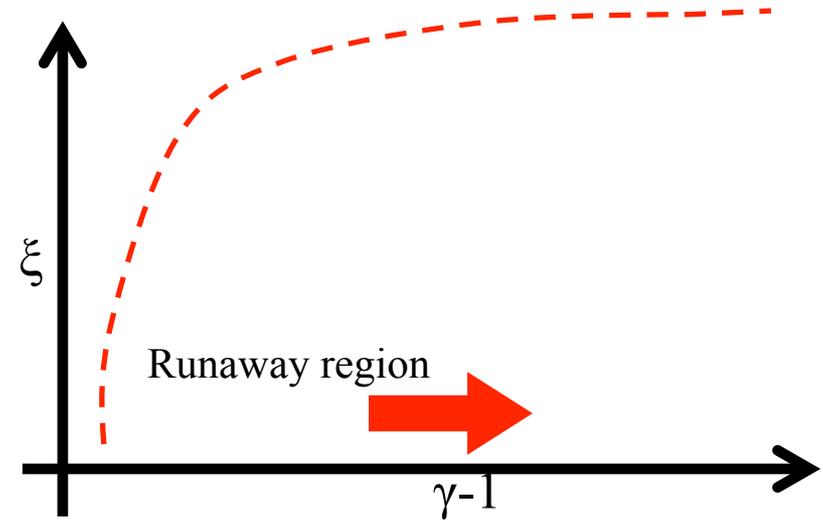
Runaway Electrons

- Runaway electrons may be present when the parallel electric field overcomes the collisional drag
 - In the absence of other physics electrons can be accelerated to arbitrarily high energies
 - Critical electric field above which electrons are accelerated given by Connor-Hastie threshold E_c
- Energy of runaway electrons thought to be limited by radiation:
 - **Bremsstrahlung radiation**: likely most important for dense plasmas with a large impurity content
 - **Synchrotron radiation**: most important mechanism in a range of parameter regimes (Guo et al. 2017)



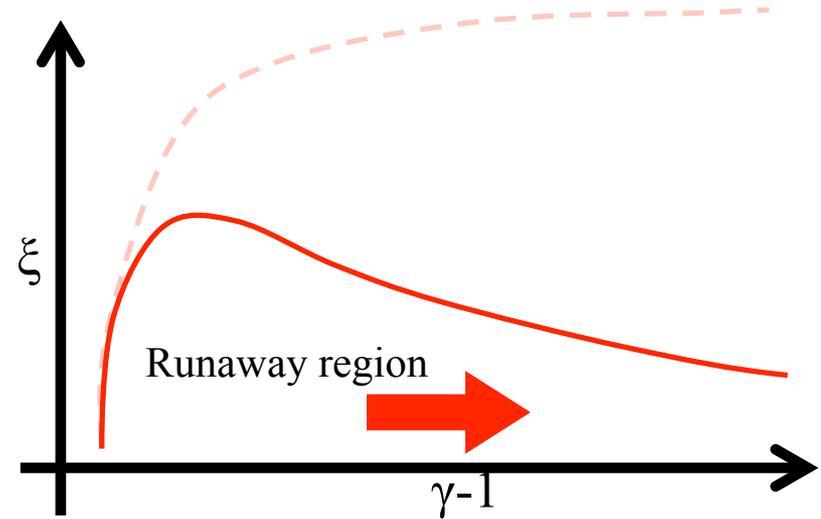
Formation of Primary Electron Tail

- For $E > E_c$ electrons in a broad region of momentum space capable of running away



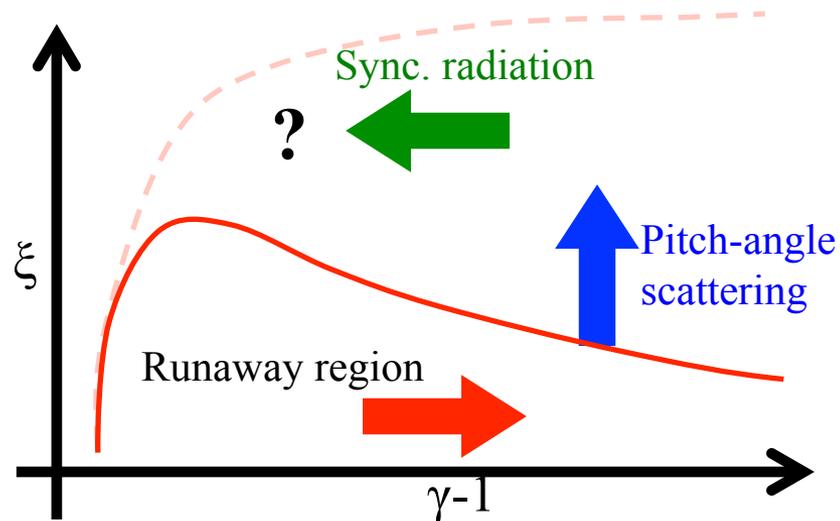
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 - Presence of synchrotron radiation constrains “runaway” region
 - **Synchrotron radius damps perpendicular energy** \Leftrightarrow narrow “channel” along $\xi=-1$ axis present



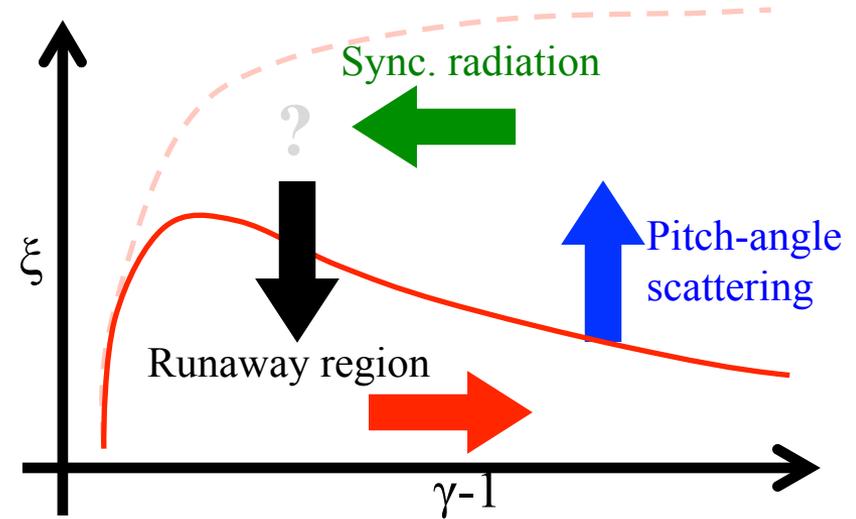
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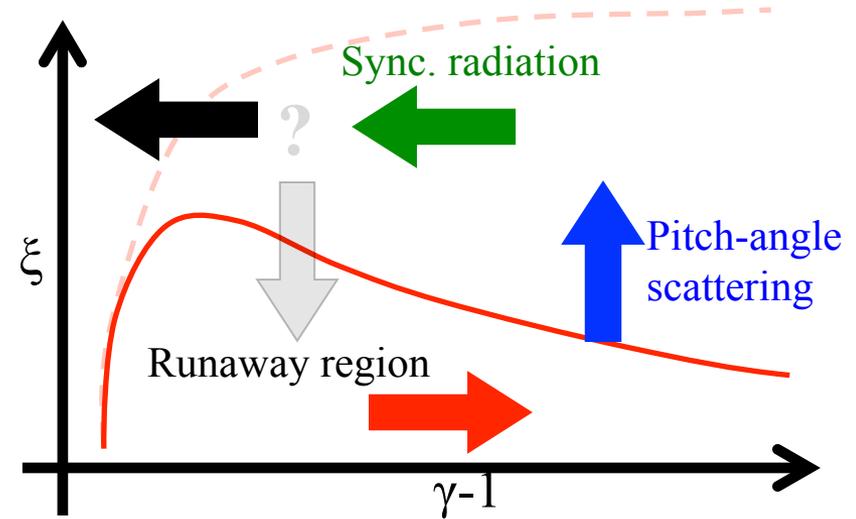
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 - Leads to confinement of runaway electron population

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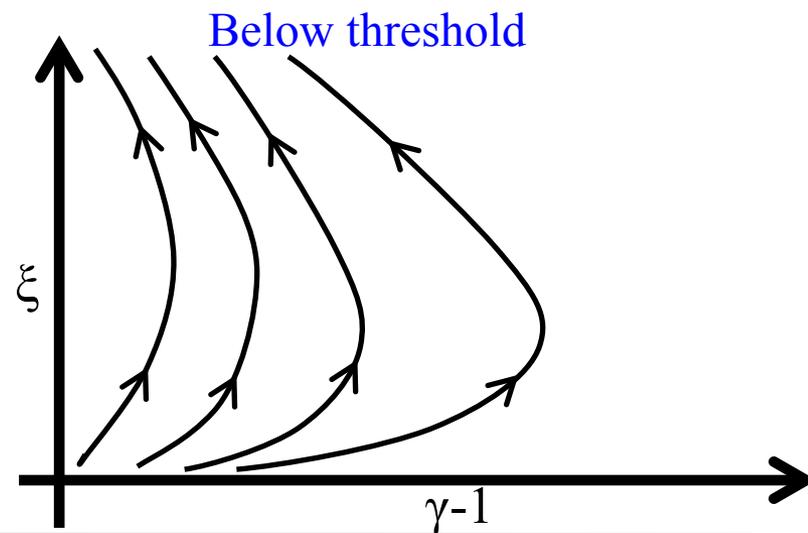
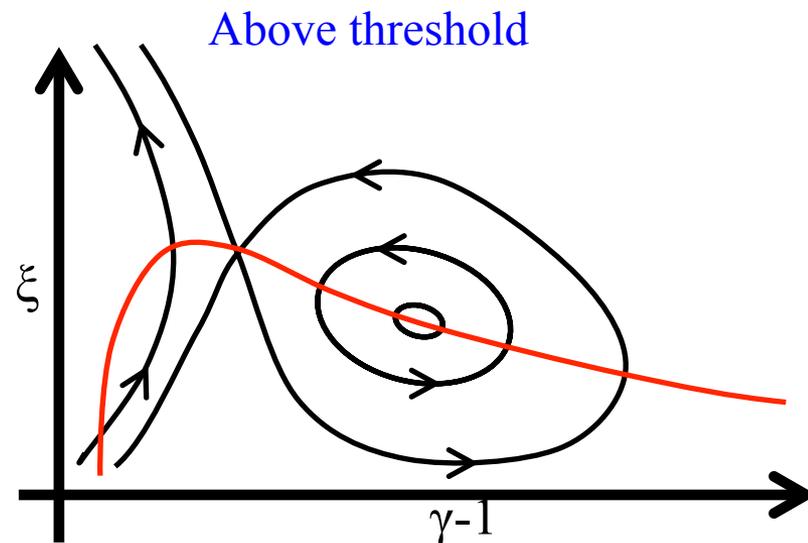


1. For a sufficiently strong electric field, primary electrons are pinched back to $\xi=-1$ axis \rightarrow runaway vortex
 - Leads to confinement of runaway electron population
2. For weaker electric fields, primary electrons return to bulk:
 - Pitch-angle scattering prevents electrons from collapsing onto $\xi=-1$ axis
 - No effective means of confining runaway electrons

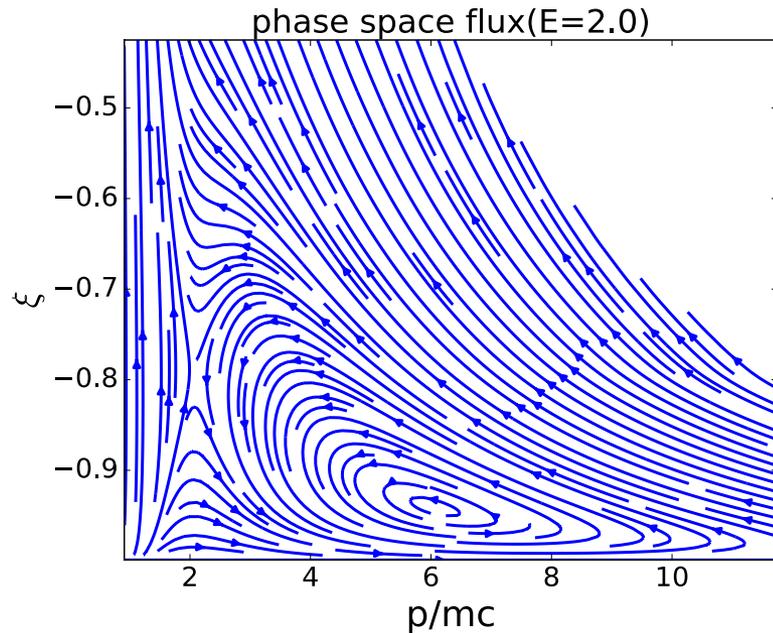
O-X Merger of Primary Distribution

- Topology of momentum space flows distinct for the two cases discussed above (Guo et al. 2017)
- Above threshold:
 - Runaway vortex present at high energy
 - Provides a means of confining runaway electrons
- Below threshold:
 - All flux lines return to electron bulk
 - No means of confining runaway electrons
- Threshold for runaway vortex distinct from Connor-Hastie threshold:

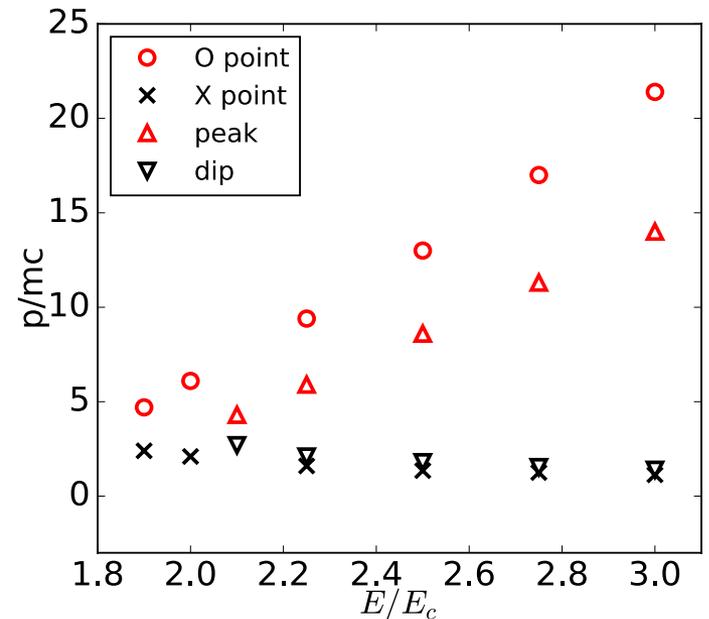
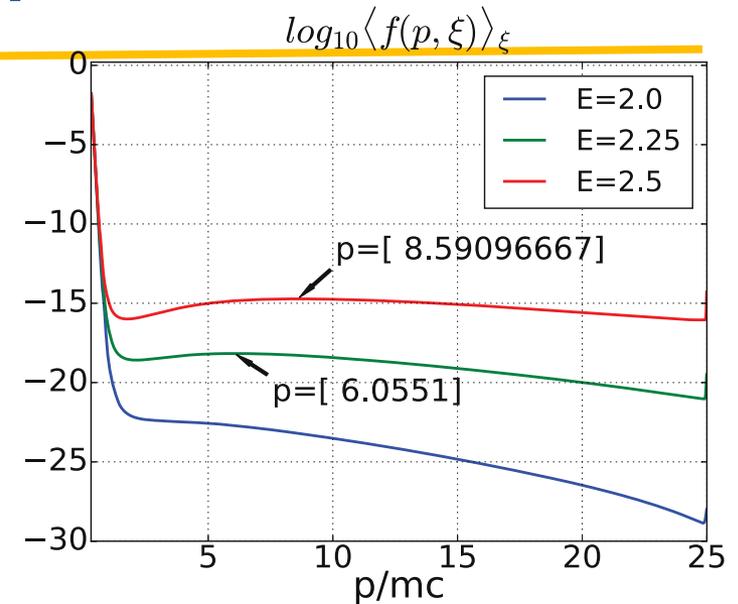
$$E_{ox} > E_c$$



Primary Runaway Distribution



- Bump and energy spread of distribution set by runaway vortex
- Threshold for bump is distinct from O-X merger threshold $E_b > E_{ox}$
 - Presence of runaway vortex does not imply a bump in runaway distribution

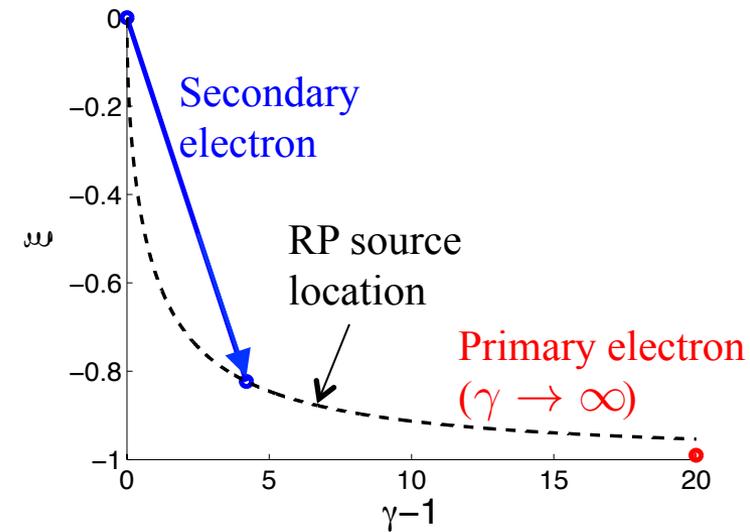


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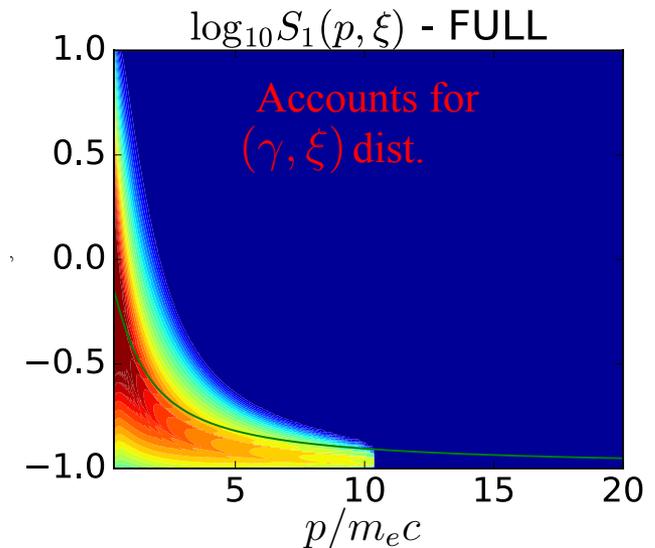
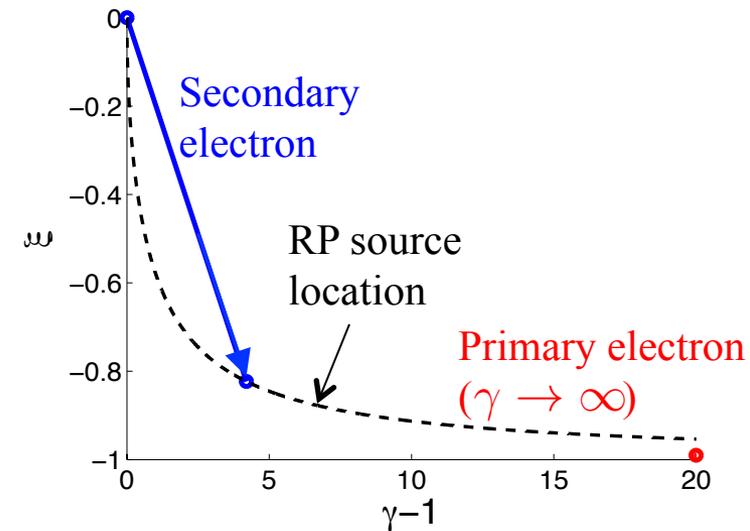
Source of Secondary Electrons → Avalanche

- Avalanche instability arises due to large-angle collisions of runaways with thermal electrons
- Rosenbluth-Putvinski (RP) secondary source provides a particularly simple picture:
 - Runaway electrons assumed to be located at infinite energy on pitch-angle axis
 - Fraction of “secondary” electrons kicked into runaway region ⇔ leads to exponential growth of runaway population



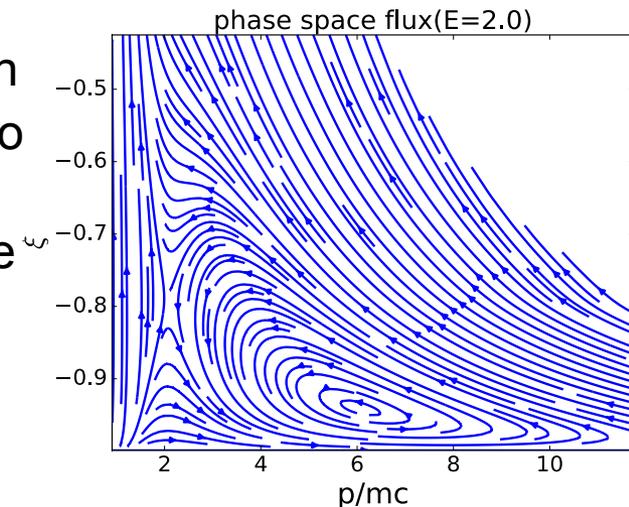
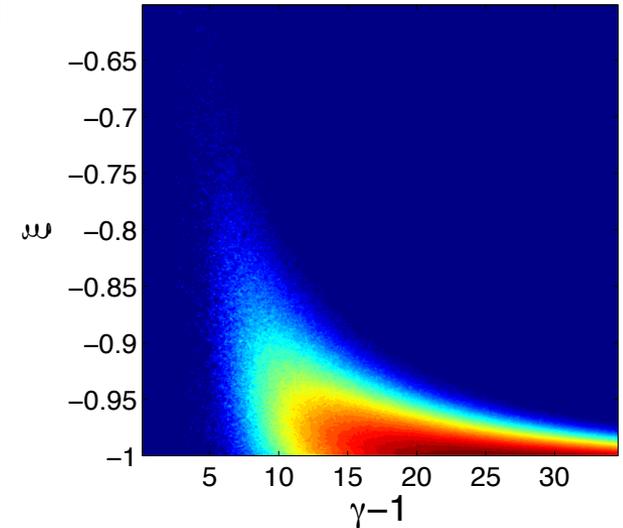
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 - Runaway electrons assumed to be located at infinite energy on pitch-angle axis
 - Fraction of “secondary” electrons kicked into runaway region ⇔ leads to exponential growth of runaway population
- A simple extension of the above model is to account for the energy/pitch-angle distribution of the primary electrons
 - Leads to a **distributed source** of secondary electrons



Runaway Electron Solvers

- The relativistic Fokker-Planck equation is solved via two distinct approaches:
- **Particle based solver**
 - Pitch-angle diffusion described by Monte Carlo collision operator [Boozer 1981], energy diffusion neglected
 - Absorbing boundary condition assumed at low energy
- **Continuum Fokker-Planck solver [Guo 2017]**
 - Considers direct solution of Fokker-Planck equation
 - Solves for tail electron distribution while matching to a Maxwellian-Jüttner distribution at low energy
 - Allows for accurate calculation of momentum space fluxes \leftrightarrow momentum space topology



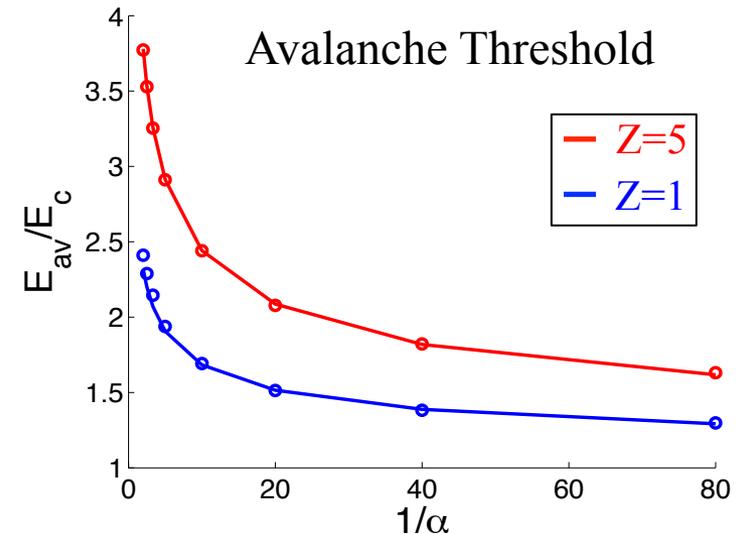
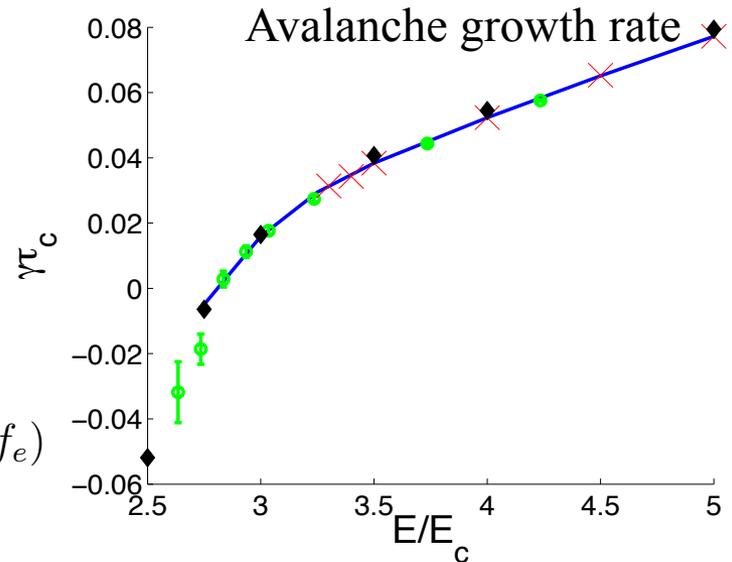
A Common Model of Avalanche Instability

- Model includes [Liu et al. 2017]:
 - Distributed secondary source term
 - Test-particle collision operator
 - Synchrotron radiation used to limit the runaway's energy

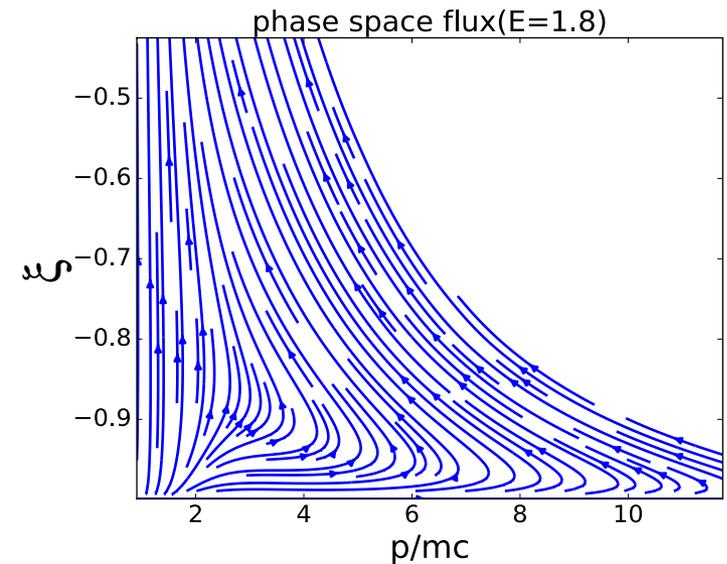
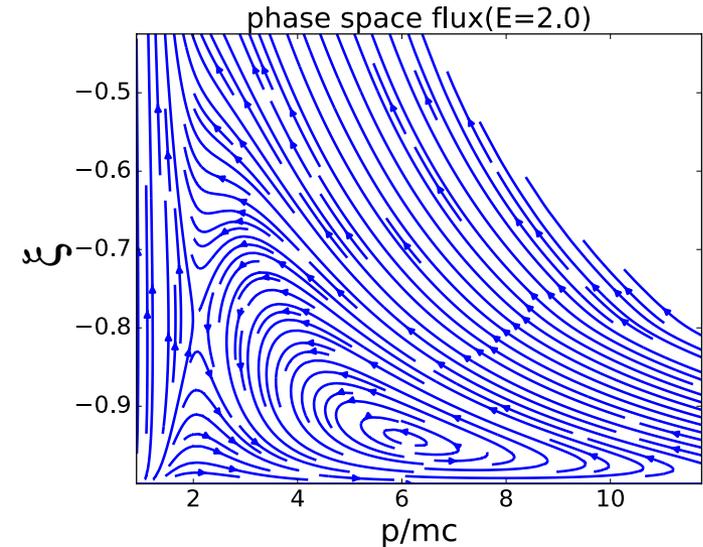
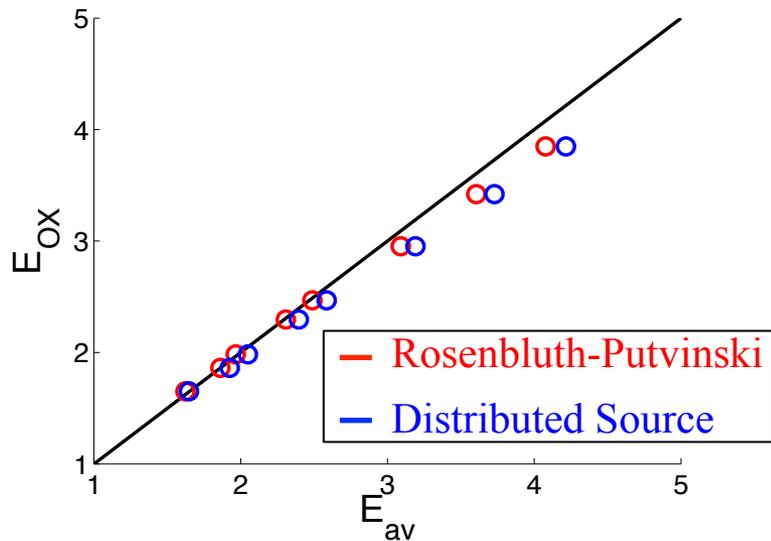
$$\frac{\partial f_e}{\partial t} - \bar{E} \left(\xi \frac{\partial f_e}{\partial p} \right) + \frac{1 - \xi^2}{p} \frac{\partial f_e}{\partial \xi} = C(f_e) + C_{rad}(f_e) + S_{sec}(f_e)$$

- Model involves several idealizations:
 - Does not incorporate self-consistent down-scattering of primary electrons
 - Uses constant Coulomb logarithm

Avalanche growth rate and threshold accurately reproduced by both the particle and continuum Fokker-Planck solvers



Relation of Avalanche Threshold to Features of Primary Distribution



- Runaway vortex \rightarrow necessary in order to confine runaway electron population at high energy
- Avalanche threshold well correlated with O-X merger over a broad range of parameters:

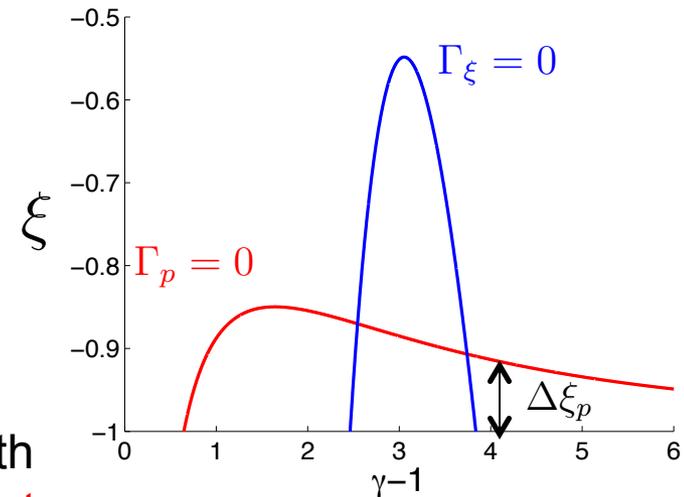
$$Z = 1 - 10, \quad \alpha = \tau_c / \tau_s = 0.05 - 0.3$$

Merger of O-X Points can be Described Analytically

- Location of O and X points set by the crossing of the $\Gamma_p = 0$ and $\Gamma_\xi = 0$ surfaces

$$\Gamma_p = [-\xi \bar{E} - C_F - \alpha p \gamma (1 - \xi^2)] f_e = 0$$

$$\Gamma_\xi = - (1 - \xi^2) \left(E - \alpha \frac{p}{\gamma} \xi + \frac{C_B}{p} \frac{\partial \ln f_e}{\partial \xi} \right) f_e = 0$$

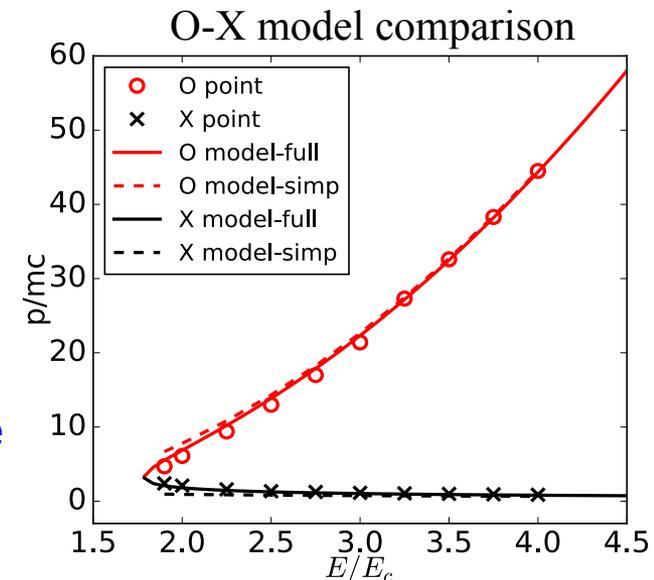


- Width of pitch-angle distribution to scale with the width of the runaway region $\Delta\xi_p$ [Decker et al. 2016, Guo et al., 2017]

$$\Delta\xi = \sqrt{2} \Delta\xi_p \approx \sqrt{2} \frac{\bar{E} - 1 - p^{-2}}{2\alpha p \gamma + \bar{E}}$$

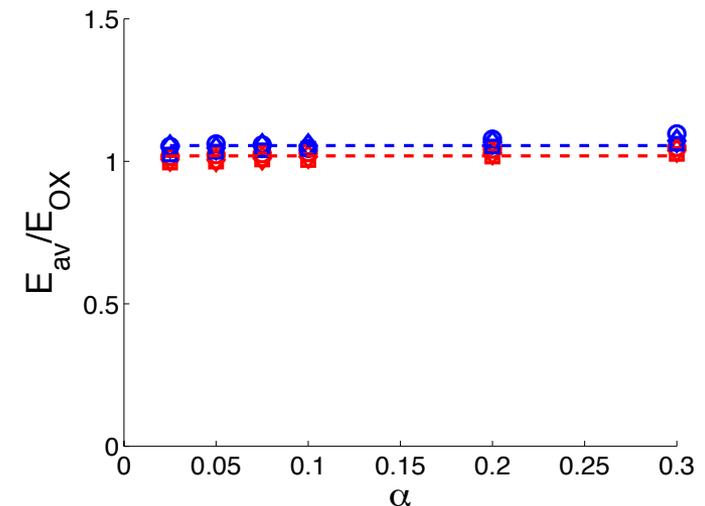
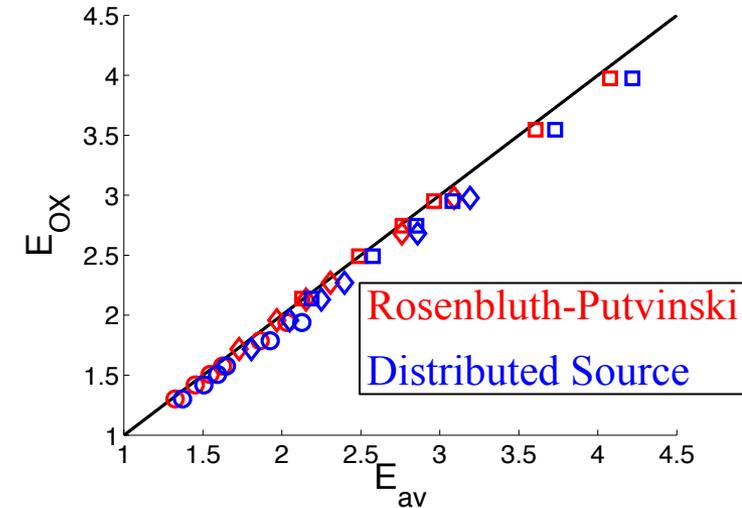
- Allows location of $\Gamma_p = \Gamma_\xi = 0$ surfaces to be estimated:

- Provides physics based means of estimating O-X merger \leftrightarrow model can be easily extended to include a variety of physical processes



Avalanche Threshold: O-X Merger

- The avalanche threshold is closely tied to the O-X merger
 - Presence of an O-point necessary to sustain a runaway population
 - Avalanche threshold is largely independent of whether Rosenbluth-Putvinski source or more comprehensive source of secondary electrons is used
- Close agreement for a range of values of synchrotron radiation/Z numbers
$$Z = 1 - 10, \alpha = \tau_c / \tau_s = 0.025 - 0.3$$



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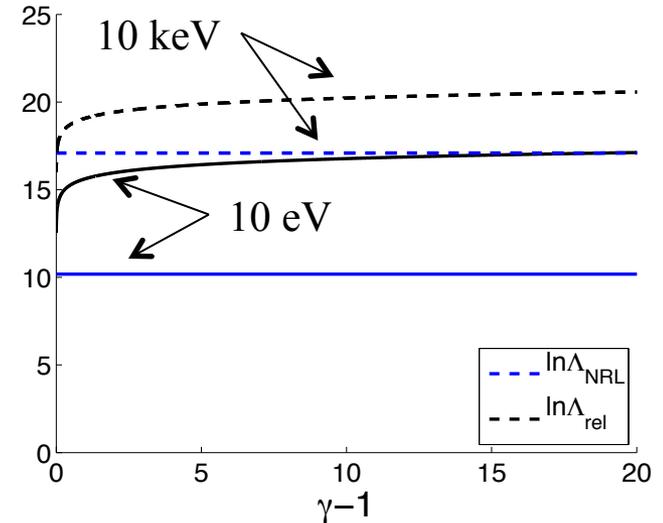
Energy Dependent Coulomb Logarithm for Runaway Electrons

- The Coulomb logarithm for relativistic electrons can deviate significantly from that of thermal electrons
- Defining $\ln \Lambda = \ln b_{max}/b_{min}$, where $b_{max} = \lambda_D$, $b_{min} = \max(b_0, \lambda_{dB})$
- However, both the classical distance of closest approach b_0 and the de Broglie wavelength λ_{dB} scale with the relative momentum between colliding electrons:

$$b_0 \sim \frac{1}{\bar{p}^2}, \quad \lambda_{dB} \sim \frac{1}{\bar{p}}, \quad \ln \Lambda_{rel} > \ln \Lambda_{Te}$$

- Thus, $b_{min} \ll b_{min}^{Te}$, $\ln \Lambda_{rel} > \ln \Lambda_{Te}$, for relativistic electrons, $\lambda_{dB}/b_0 \approx 137/2$ hence [Solodov-Betti 2009]:

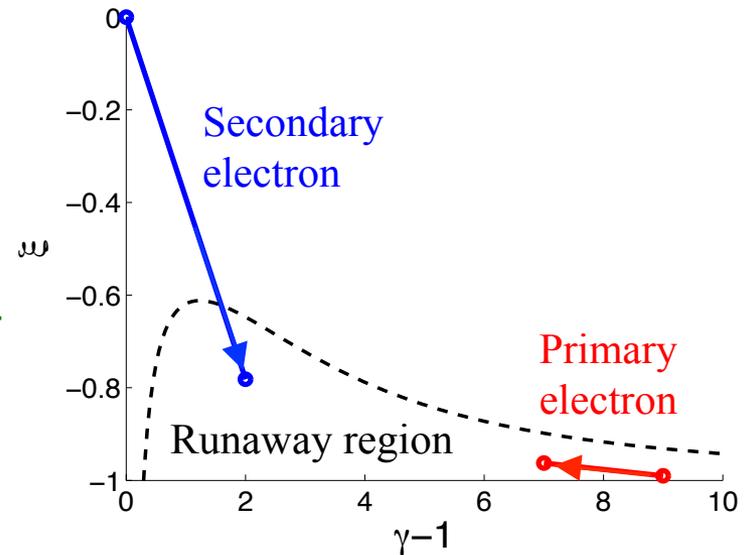
$$\ln \Lambda_{rel}(p) = \ln \frac{\lambda_D}{\lambda_{dB}(p)}$$



- Form described in [Mosher 1975, Martin-Solis 2017] for free-free collisions use $b_{min} = b_0 \rightarrow$ leads to significantly larger Coulomb logarithm

Conservative Large-Angle Collision Operator

- Large-angle collisions often incorporated via a knock-on “source”
 - Primaries evolved by Fokker-Planck collision operator
 - Leads to inconsistent formulation \Leftrightarrow large-angle scattering of primary electrons not accounted for



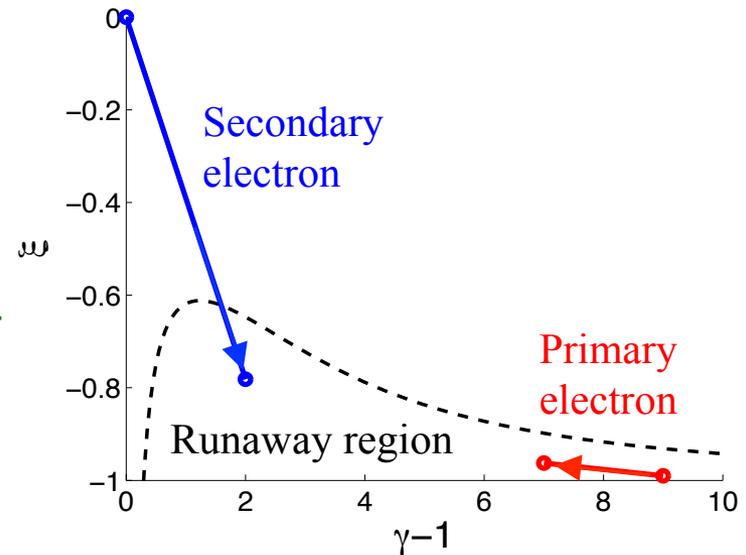
- Redistribution of primaries can be accounted for by considering complete form of linearized Boltzmann collision operator:

$$\left. \frac{\partial f_e}{\partial t} \right|_{LA} = \frac{n_b}{2\pi p^2} \int d^3 p' v' \frac{d\sigma_M(p', p)}{dp} \Pi(p', \xi', p; \xi) f_1(\mathbf{p}') - \frac{1}{2} n_b \delta(\mathbf{p}) \int d^3 p_1 v_1 \sigma_M(p_1) f_1(\mathbf{p}_1) - \frac{1}{2} n_b v \sigma_M(p) f_1(\mathbf{p})$$

- Ensures energy and momentum conservation from large-angle collisions

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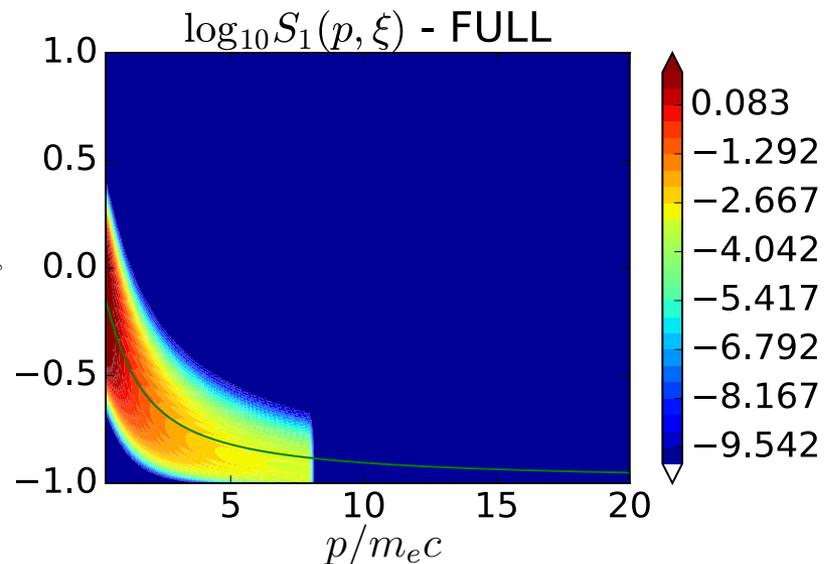
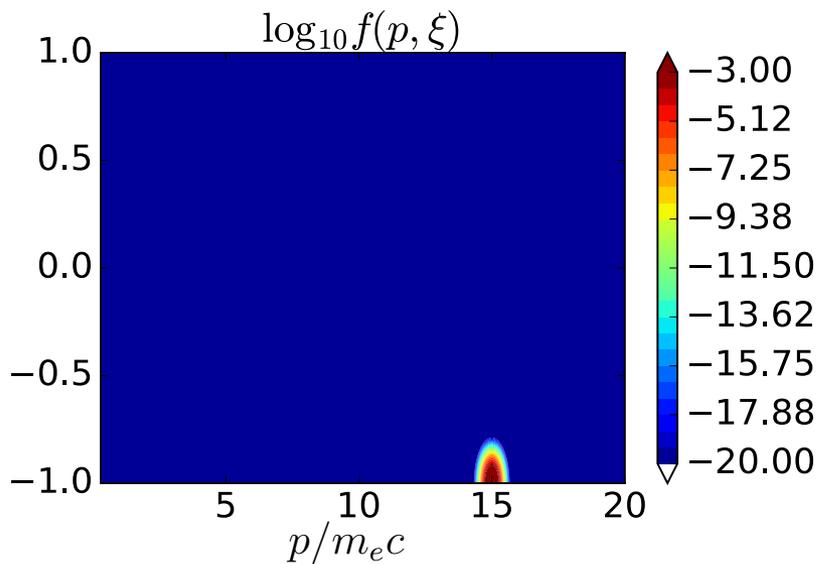
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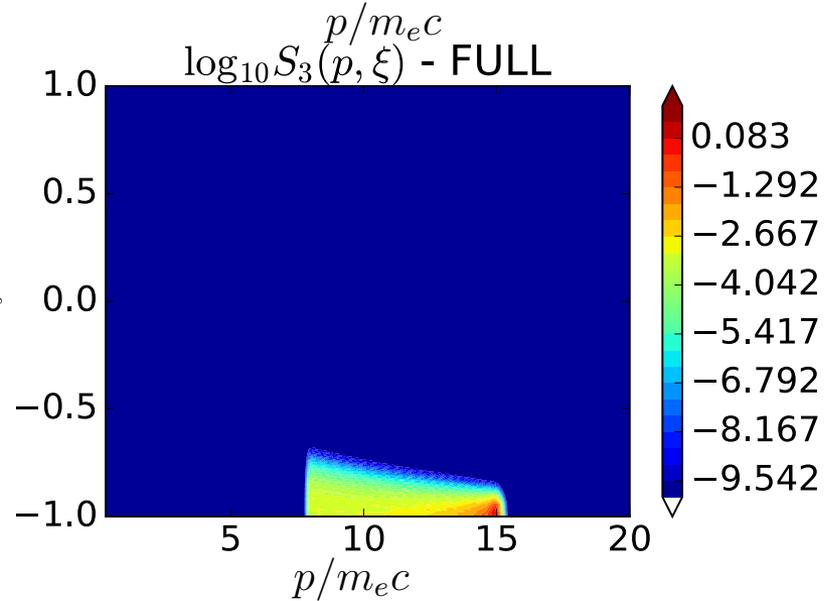
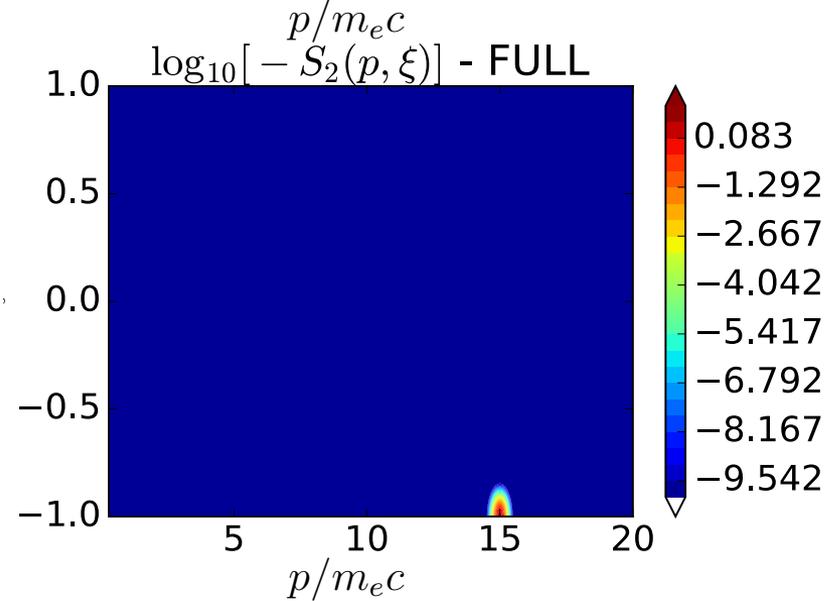
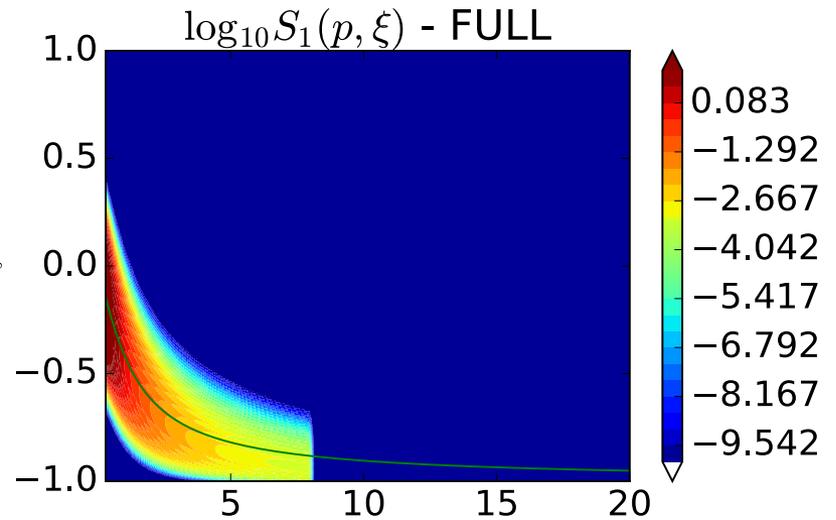
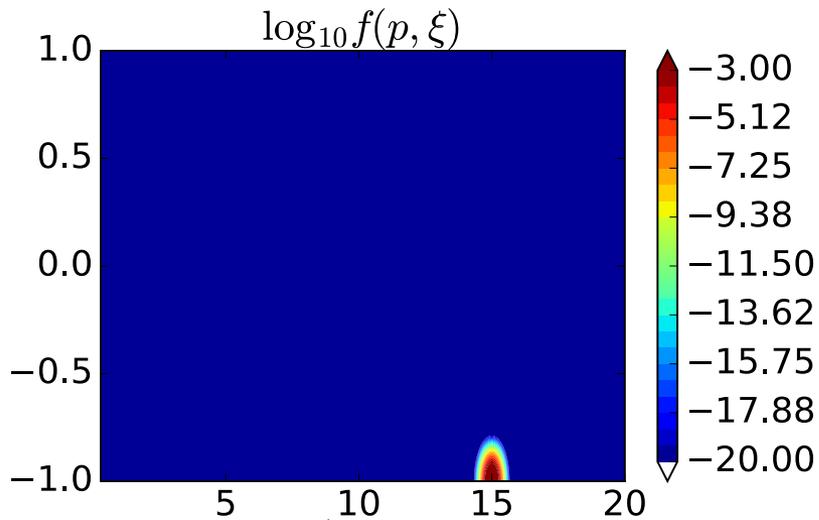
Neglected here

- Ensures energy and momentum conservation from large-angle collisions

Back Reaction onto Primary Electrons

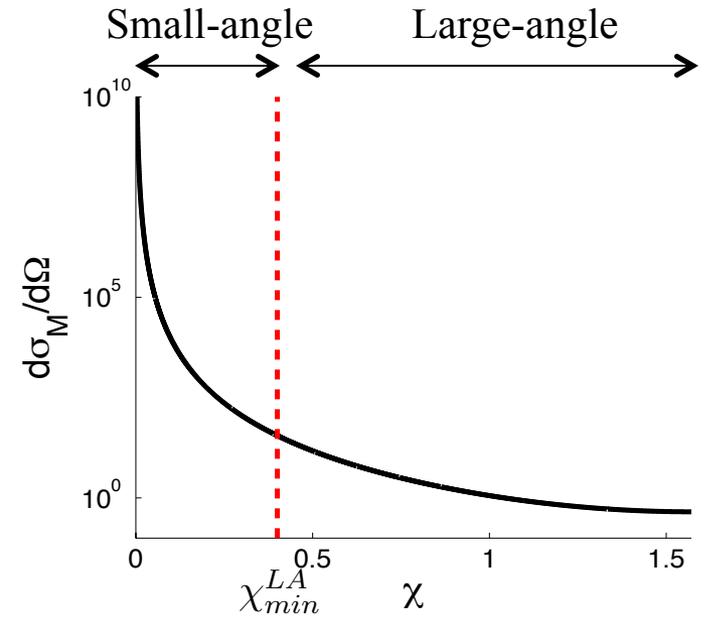


Back Reaction onto Primary Electrons



Avoiding Double Counting Collisions

- Addition of down scattering terms can result in double counting collisions



Avoiding Double Counting Collisions

- Addition of down scattering terms can result in double counting collisions
- The choice of an intermediate cutoff angle can be constrained by noting:
 1. The large-angle collision operator assumes the thermal electron population to be cold:

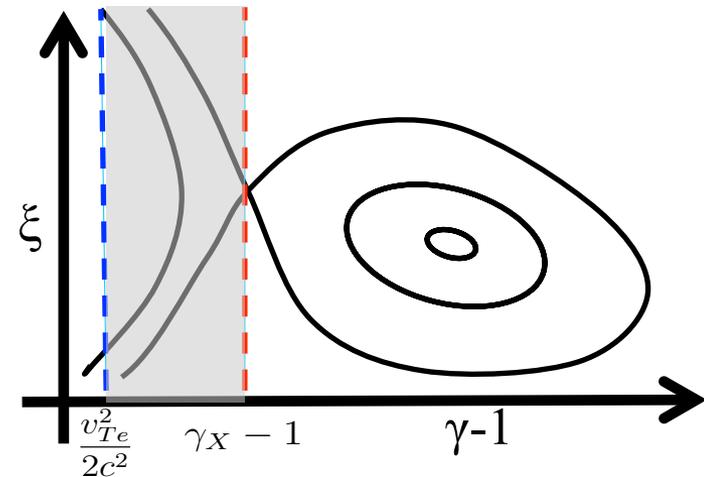
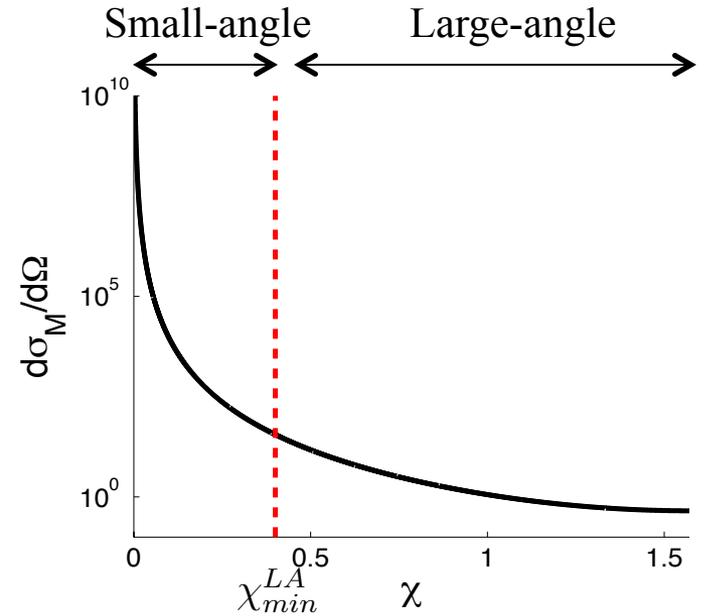
$$m_e c^2 (\gamma_{min} - 1) \gg m_e v_{Te}^2 / 2$$

2. Only secondary electrons whose energy is greater/comparable to the X point location can runaway:

$$\gamma_{min} - 1 < \gamma_X - 1$$

- Center-of-mass scattering angle χ_{min}^{LA} related to γ_{min} by:

$$\sin^2 \frac{\chi_{min}^{LA}}{2} = \frac{\gamma_{min} - 1}{\gamma - 1}$$



Modified Coulomb Logarithm

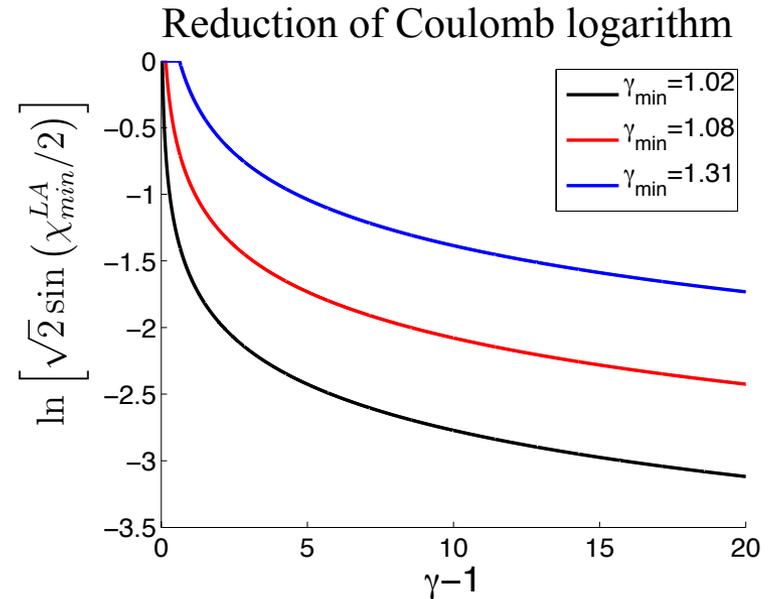
- Introduction of additional cutoff angle modifies Coulomb logarithm used in Fokker-Planck equation
- The drag coefficient in the Fokker-Planck equation is proportional to:

$$\frac{\langle \Delta p \rangle}{\Delta t} \sim \int_{\chi_{min}}^{\chi_{min}^{LA}} d\chi \sin \chi \sin^2 \frac{\chi}{2} \frac{d\sigma_m}{d\Omega} \sim \ln \left[\frac{2}{\chi_{min}} \sin \left(\chi_{min}^{LA}/2 \right) \right] + \mathcal{O}(1)$$

- Coulomb logarithm reduced by a factor:

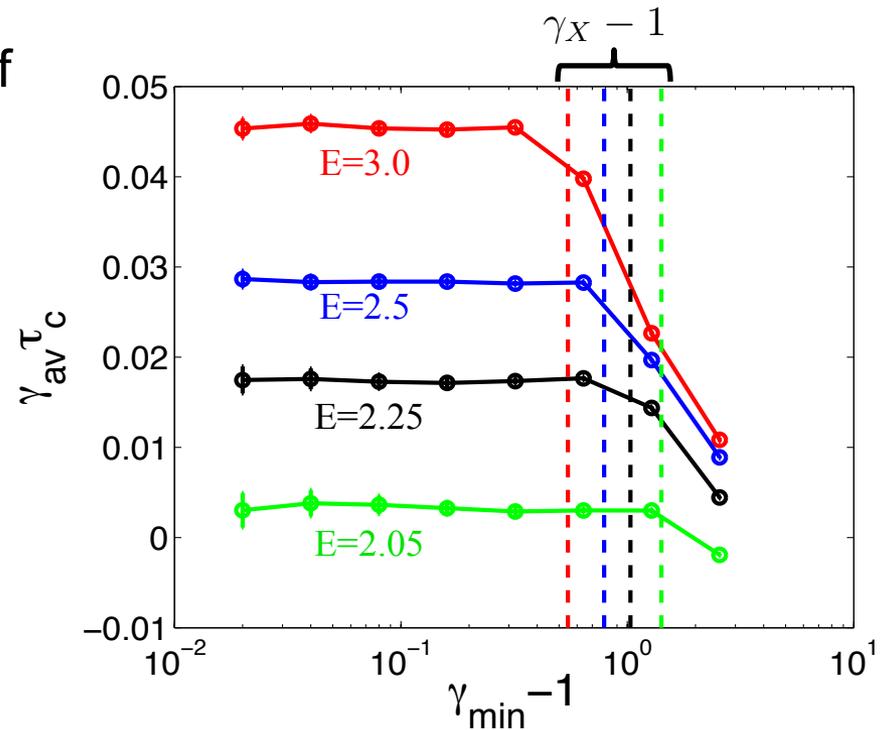
$$\ln \left[\sqrt{2} \sin \left(\chi_{min}^{LA}/2 \right) \right] \leq 0$$

- Leads to a significant reduction of Coulomb logarithm used in Fokker-Planck equation (10-20%)



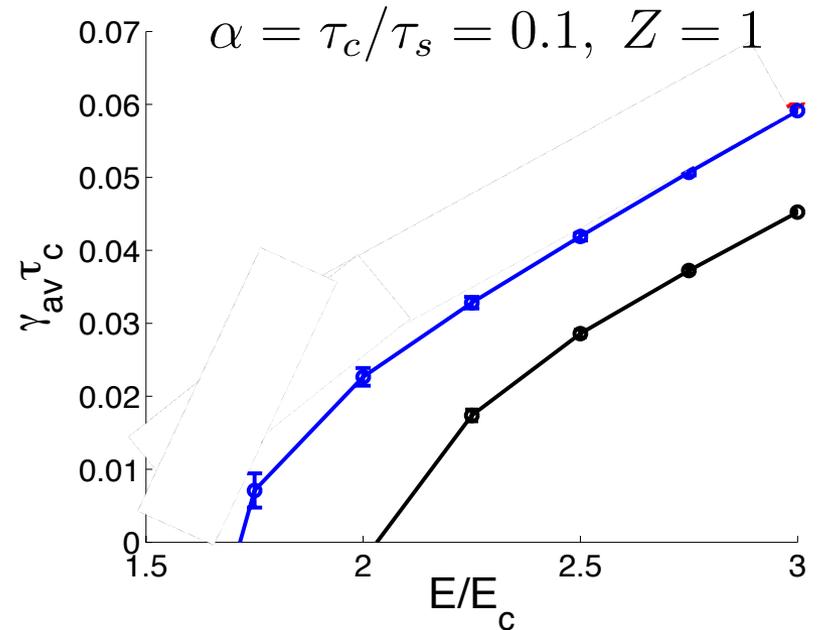
Sensitivity to Intermediate Cutoff

- Necessary to assess the sensitivity of the avalanche growth rate to the intermediate cutoff angle
[$\chi_{min}^{LA} = \chi_{min}^{LA}(\gamma_{min})$]
- Broad regime of insensitivity evident:
 - For γ_{min} too large, the avalanche growth rate underestimated \Leftrightarrow incorrectly omits secondary electrons
 - Formalism invalid for energies comparable to the thermal energy



Contrasting Current Model with Previous Model

- Addition of conservative large-angle collision operator + modified Coulomb logarithm reduces avalanche growth rate

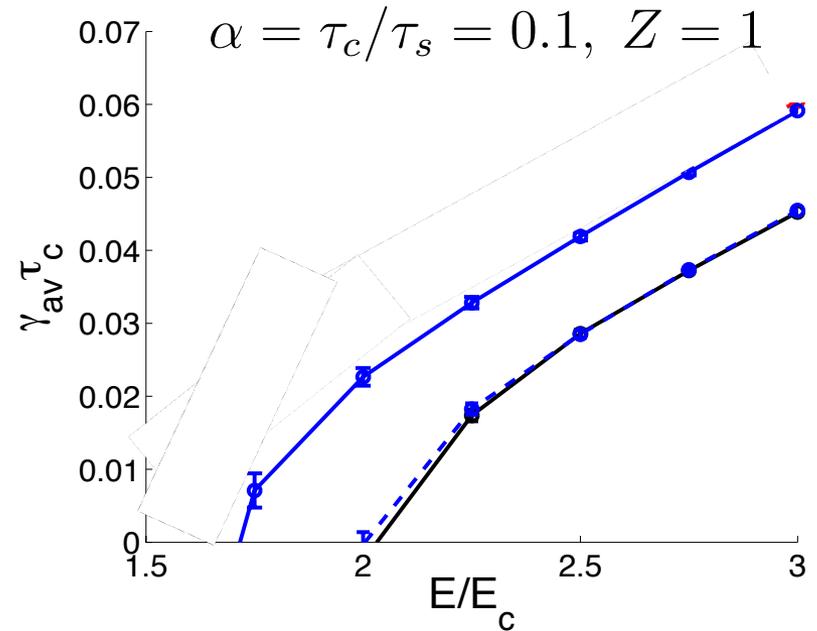


— Secondary Source + $\ln \Lambda = \text{const}$

— Boltz. + $\ln \Lambda_{rel} - \left| \ln \left[\sqrt{2} \sin(\chi_{min}^{LA}/2) \right] \right|$

Contrasting Current Model with Previous Model

- Addition of **conservative large-angle collision operator + modified Coulomb logarithm** reduces avalanche growth rate
- Useful to identify which additional physics has the strongest impact on avalanche growth rate:
 - Consider additional case with $\ln \Lambda = \ln \Lambda_{rel}(p)$, but without conservative large-angle collision operator
 - **Conservative large-angle collision operator has a modest impact on avalanche growth rate once double counting constraint is accounted for!**



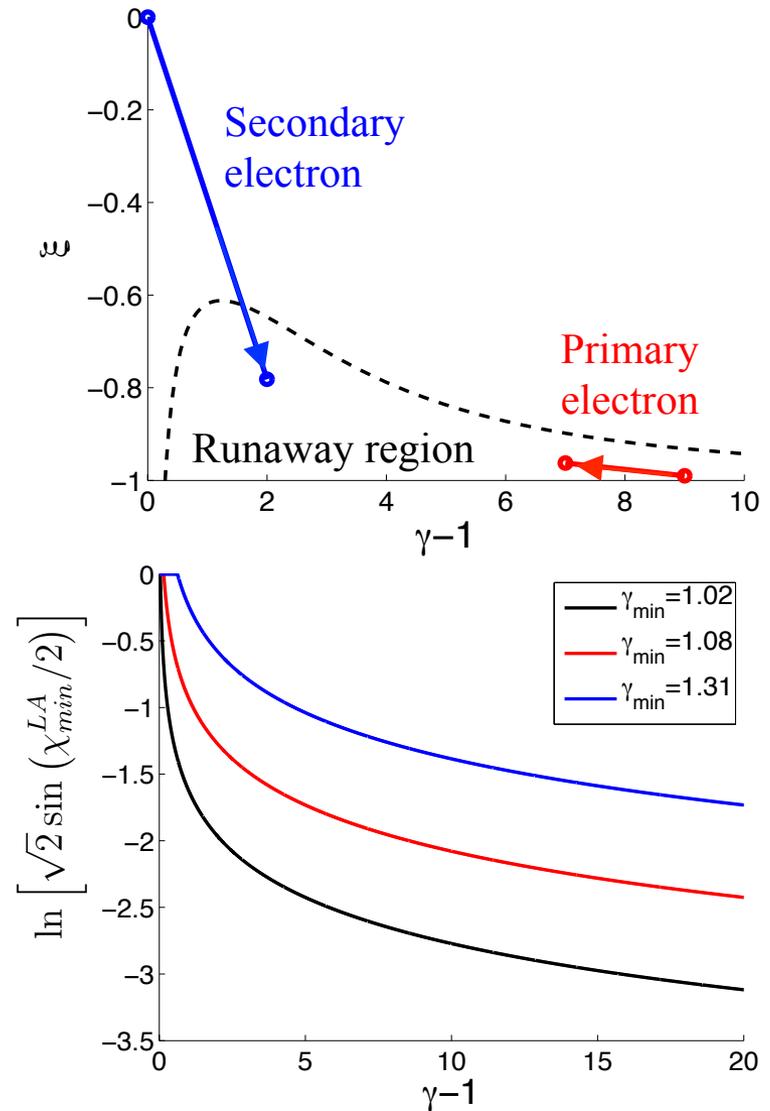
- Secondary Source + $\ln \Lambda = \text{const}$
- Boltz. + $\ln \Lambda_{rel} - \left| \ln \left[\sqrt{2} \sin(\chi_{min}^{LA}/2) \right] \right|$
- Secondary Source + $\ln \Lambda_{rel}(p)$

Impact of Large-Angle vs. Small-Angle Collisions

- Use of conservative large-angle collision operator involves a tradeoff:
 - Introduces additional down-scattering of primary electrons
 - To avoid double counting, Coulomb logarithm is reduced by:

$$\ln \Lambda = \ln \Lambda_{rel} - \left| \ln \left[\sqrt{2} \sin \left(\frac{\chi_{min}^{LA}}{2} \right) \right] \right|$$

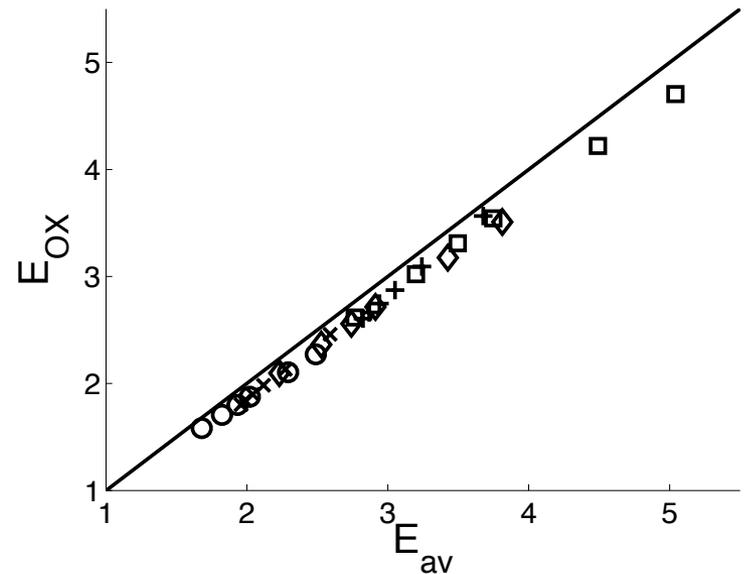
- Both modifications scale with $1 / \ln \Lambda$
 - Leads to modest impact on avalanche growth rate



Avalanche Threshold Well Described by O-X Merger

- O-X merger model of primary distribution can be modified to incorporate a more accurate Coulomb logarithm
- Predicted avalanche threshold in good agreement with O-X merger for a broad parameter range

$$Z = 1 - 10, \quad \alpha = \tau_c / \tau_s = 0.025 - 0.3,$$
$$\ln \Lambda_{NRL} = 5 - 20$$



$$\text{Boltz.} + \ln \Lambda_{rel} - \left| \ln \left[\sqrt{2} \sin \left(\chi_{min}^{LA} / 2 \right) \right] \right|$$

Conclusions

- O-X merger of primary distribution identified as a reliable indicator of avalanche instability \Leftrightarrow runaway vortex necessary to confine runaway electron population:
 - A robust, and highly malleable, analytic model describing the O-X merger of the primary distribution has been developed
- A conservative large-angle collision operator appropriate to the description of runaway electrons has been derived:
 - Systematic procedure for delineating small and large angle collision operators developed
- A more comprehensive avalanche model that self-consistently treats large-angle scattering events and incorporates a Coulomb logarithm appropriate for runaway electrons has been derived
 - Significantly modifies the avalanche growth rate and threshold



Relation of Avalanche Threshold to Features of Primary Distribution

- Interested in identifying how the avalanche threshold is correlated with features of the **primary** distribution function
- We will consider a comparison with two features of the primary distribution:
 1. Merger of O-X point \Leftrightarrow necessary in order to confine runaway electron population at high energy [Guo et al. 2017]
 2. Formation of a pitch-angle averaged bump \Leftrightarrow prominent role in attractor picture [Aleynikov-Breizman 2014]
- Avalanche threshold most closely correlated with O-X merger over a broad range of parameters:

$$Z = 1 - 10, \quad \alpha = \tau_c / \tau_s = 0.05 - 0.3$$

