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3D non-linear MHD modelling of Massive Gas Injection (MGI) -triggered disruptions in JET

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# **Outline**

### The model

- Description of a typical simulation
- Mechanisms leading to the Thermal Quench (TQ)
- Physics of the TQ and I<sub>p</sub> spike

# The model

## The JOREK model for MMI (i.e. MGI or SPI) simulations

Reduced MHD, no diamagnetic effects

8 variables:

- Poloidal magnetic flux  $\,\psi\,$  (  ${f B}=F_0
  abla \phi+
  abla \psi imes
  abla \phi$  )
- Toroidal current density  $j = \Delta^* \psi$
- Electric/flow potential u■ Parallel velocity  $v_{\parallel}$  →  $\mathbf{v} = v_{\parallel} \mathbf{B} R^2 \nabla u \times \nabla \phi$ (assumed common for all species)
- Vorticity  $\omega = \frac{1}{R} \frac{\partial}{\partial R} \left( R \frac{\partial u}{\partial R} \right) + \frac{\partial^2 u}{\partial Z^2}$
- Impurity mass density  $ho_{imp}$  (summed over all charge states)

- Total mass density (main ions + impurities)  $\rho$ 

- **—** Temperature T, assumed common to all species
- Coronal equilibrium (CE) assumption provides charge state distribution, radiation losses, ionization energy (ADAS data)
  - Benchmark with M3D-C1 and NIMROD suggests that CE assumption slows down the radiative collapse

$$\begin{split} \text{Ohm's law} \quad & \frac{\partial \psi}{\partial t} = \eta \left( T_e \right) \Delta^* \psi - R \left\{ u, \psi \right\} - F_0 \frac{\partial u}{\partial \phi} \qquad (\texttt{+ hyper-resistivity}) \\ \text{Vorticity} \quad & R \nabla \cdot \left[ R^2 \left( \rho \nabla_{pol} \frac{\partial u}{\partial t} + \nabla_{pol} u \frac{\partial \rho}{\partial t} \right) \right] = \frac{1}{2} \left\{ R^2 \left| \nabla_{pol} u \right|^2, R^2 \rho \right\} + \left\{ R^4 \rho \omega, u \right\} \\ & - R \nabla \cdot \left[ R^2 \nabla_{pol} u \nabla \cdot (\rho \mathbf{v}) \right] + \left\{ \psi, j \right\} - \frac{F_0}{R} \frac{\partial j}{\partial \phi} \\ & + \left\{ P, R^2 \right\} + R \mu \left( T_e \right) \nabla_{pol}^2 \omega. \quad (\texttt{+ hyper-viscosity}) \end{split}$$

$$\text{// momentum} \quad B^{2} \frac{\partial}{\partial t} \left( \rho v_{\parallel} \right) = -\frac{1}{2} \rho \frac{F_{0}}{R^{2}} \frac{\partial}{\partial \phi} \left( v_{\parallel} B \right)^{2} - \frac{\rho}{2R} \left\{ B^{2} v_{\parallel}^{2}, \psi \right\} - \frac{F_{0}}{R^{2}} \frac{\partial P}{\partial \phi} + \frac{1}{R} \left\{ \psi, P \right\} \\ -B^{2} \nabla \cdot \left( \rho \mathbf{v} \right) v_{\parallel} + B^{2} \mu_{\parallel} \left( T_{e} \right) \nabla_{pol}^{2} v_{\parallel}.$$

$$\begin{array}{ll} \text{Mass conservation} \\ \text{(impurities)} \end{array} & \quad \frac{\partial}{\partial t} \rho_{imp} = -\nabla \cdot \left( \rho_{imp} \mathbf{v} \right) + \nabla \cdot \left( D_{imp} \nabla \rho_{imp} \right) + \underbrace{S_{imp}} \text{MGI/SPI} \end{array}$$

(all)  $\frac{\partial}{\partial t}\rho = -\nabla \cdot (\rho \mathbf{v}) + \nabla \cdot [D_D \nabla (\rho - \rho_{imp})] + \nabla \cdot (D_{imp} \nabla \rho_{imp}) + S_D + S_{imp}$ 

Energy density 
$$\frac{\partial P^*}{\partial t} + \nabla \cdot (\mathbf{v}P^*) = -(\gamma - 1) P \nabla \cdot \mathbf{v} + \frac{2}{3R^2} \eta (T_e) j^2 + \nabla \cdot \left(\kappa_{\perp} \nabla_{\perp} T + \kappa_{\parallel} \nabla_{\parallel} T\right)$$
  
where  $P^* \equiv P + (2/3)n_{imp}E_{ion}$   
 $-n_e n_{imp}P_{rad} (T_e) + \frac{\gamma - 1}{2} \mathbf{v} \cdot \mathbf{v} (S_D + S_{imp}).$ 

# Description of a typical simulation



Simulated pulse: JET #85943

- Ohmic, 2MA, 3T, T<sub>e0</sub>=3.3keV, n<sub>e0</sub>=2.1x10<sup>19</sup>m<sup>-3</sup>
- Pure Ar MGI from Disruption Mitigation Valve 1 (DMV1) at 33 bar into a healthy plasma
- Simulation setup:
  - Time dependence of Ar source based on Euler equations, but P<sub>DMV</sub> is reduced to account for fuelling efficiency <<100%</p>
  - Realistic (Spitzer) resistivity and (turbuent) viscosity & diffusivities (+ scans)
  - For numerical reasons, parallel flows are artificially damped







Initial state

(Poloidal cuts in the plane of DMV1)

#### 2.70ms after gas arrival



Thin radiating ring → cold front → current profile contraction
 Growth of tearing modes (here m/n=3/1 island visible)

#### 3.77ms after gas arrival



Growth of 2/1 modeSome stochasticity

#### 3.86ms after gas arrival



- Thermal Quench (TQ) onset
- Stochastic region expands fast
- Heat flux conducted into region where  $n_{imp}$  is large  $\rightarrow$  large localized  $P_{rad}$  11

#### 4.08ms after gas arrival



Global stochasticity  $\rightarrow$  global T<sub>e</sub> flattening (TQ)

#### 4.54ms after gas arrival



After the TQ, flux surfaces start reappearing in the core

# Mechanisms leading to the Thermal Quench (TQ)



## A realistic wall is needed to reproduce the pre-TQ $I_p$ evolution



Using a realistic resistive wall (red and magenta curves), the pre-TQ I<sub>p</sub> drop is well matched

- Pre-TQ  $I_p$  drop ~independent of the gas amount (red vs. magenta)

- In contrast, an ideal wall close to the plasma (black curve) makes the pre-TQ I<sub>p</sub> drop too large
- Consistent with theory [Artola et al., submitted]
  - Current lost in the edge is largely re-induced in the still hot plasma
  - The process does not depend on the timescale, only on geometry 16

# Physics of the TQ and I<sub>p</sub> spike

## Context

Mechanism of the I<sub>p</sub> spike [D. Biskamp, *Nonlinear MHD*]:

- $\blacksquare$  MHD relaxation at TQ  $\rightarrow$  broadening of current profile
  - Detailed mechanism, according to A. Boozer [PPCF 2018 and NF 2019]: magnetic stochasticity connects regions with different  $j_{\parallel}/B$  $\rightarrow$  excitation of shear Alfvén waves by  $\nabla_{\parallel}(j_{\parallel}/B)$  term in vorticity equation  $\rightarrow$  redistribution (homogenization) of  $j_{\parallel}/B$
- equation  $\rightarrow$  redistribution (homogenization) of  $j_{\parallel}/B$ **Conservation of magnetic helicity**  $H \equiv \int \mathbf{A} \cdot \mathbf{B} \, dV \rightarrow \mathbf{I}_p$  has to increase
- Can be modelled in 2D via hyper-resistivity (mean field model)
  - Done in JOREK by Javier Artola
- In the past, 3D non-linear MHD simulations always underestimated the Ip spike (as far as I know)
  - $\Rightarrow$  MHD relaxation not well captured?
  - ⇒ Unreliable predictions on electron stochastic losses (which play an important role in runaway electron generation)?



This year, got first JOREK simulations with an I<sub>p</sub> spike comparable to experimental data

- Even larger in certain cases!
- Large I<sub>p</sub> spike associated to violent MHD activity and small scale excitation across the whole plasma (see simulation in next slide)
  - Could well be Alfvén waves turbulence



## What determines the I<sub>p</sub> spike height?

Unfortunately, no clear-cut conclusions yet...

- Simulations take weeks and many fail during the TQ
- However, based on existing simulations, my *impression* is that:
  - The I<sub>p</sub> spike height correlates with the amplitude of the low n modes (m/n=2/1, 3/2, ...)
  - ...which correlates with the sharpness of the skin current generated by the cold front
  - ...which correlates with
    - the Lundquist number used in the simulation
    - the « abruptness » of the radiative collapse
      - In fact, JOREK simulations with the largest I<sub>p</sub> spikes had a bug in the call to ADAS routines, making the radiative cooling rate artificially large
      - Now that the bug is solved, the I<sub>p</sub> spike is smaller than in the experiment again (but the radiative collapse may be too slow due to the coronal eq. assumption...)

#### **Testing Boozer's formula**

Boozer provides a « heuristic » formula connecting the hyper-resistivity  $(\Lambda_m)$  of the 2D mean field model to the field lines diffusion coefficient  $(D_{FL} \sim \Psi_t^2/N_t)$  of the 3D system [A. Boozer, PPCF 61 (2018) 024002]:

$$\Lambda_m \approx \frac{1}{144} \frac{2\kappa_0}{1+\kappa_0^2} \frac{\mu_0}{4\pi} \frac{V_A^2 \Psi_t^2}{N_t}$$

 $(\Psi_t \equiv \text{radial extent of stochastic region in toroidal flux units})$  $N_t \equiv \text{number of toroidal turns for a field line to travel across this region})$ 

- Analogous to Rechester-Rosenbluth formula but for magnetic helicity diffusion instead of heat diffusion
- Could allow extracting an estimate of D<sub>FL</sub> from I<sub>p</sub>(t) experimental data, with no need for 3D simulations!
- $\rightarrow$  Use JOREK simulations to test Boozer's formula

- Idea: run 2D « mean field model » simulations and look for hyperresistivity settings which allow matching 3D simulations
- Parameterization of hyper-resistivity:  $\Lambda_m = \Lambda_{m0}(1 + \tanh((\Psi_N \Psi_{N,cut})/0.01))/2$ In principle,  $\Psi_{N,cut} \leftrightarrow$  edge of the stochastic region

Good match to 3D simulation for  $\Psi_{N,cut} = 0.87$  and  $\Lambda_{m0} = 2.5 \times 10^{-6}$ 



Boozer's formula converted into JOREK units reads:  $\Lambda_{m0} = 2 \times 10^{-5} / N_t$ 

 $\Rightarrow$  Estimate N<sub>t</sub> from field line tracing in 3D simulation

Crude method: initialize many field lines at  $\Psi_N^{1/2}=0.4$ , track them and plot their radial position vs. number of turns (here at t=4.654ms)



 $\Rightarrow$  By eye, N<sub>t</sub> ~ 3  $\rightarrow$   $\Lambda_{m0}$  ~ 7x10^{-6}, to be compared to 2.5x10^{-6} found in previous slide

 $\Rightarrow$  Order of magnitude seems OK

# Conclusion

## Conclusion

JOREK simulations suggest the following picture for MGI-triggered disruptions:



(Note: role of the 1/1 mode still to be clarified)

- The above picture seems to apply (to some extent) to SPI as well
- The height of the I<sub>p</sub> spike gives information on the « intensity » of the whole process

Too small a spike likely indicates underestimated stochasticity

Preliminary investigations of Boozer's formula are promising





## **Interferometry comparison**



# About the choice of resistivity $\eta_0$ and hyper-resistivity $\eta_H$ $\partial_t \Psi = \dots + \eta_0 (T/T_0)^{-3/2} j_{\phi} + \eta_H \Delta j_{\phi}$ For numerical stability

Current density profile (axisymmetric Ar MGI simulations)



- Larger  $\eta_0 \rightarrow$  smoother skin current  $\rightarrow$  milder MHD and TQ
- ⇒ Need to use a realistic  $\eta_0$ , at least in the pre-TQ phase
- η<sub>H</sub>=10<sup>-10</sup> smooths profile w/o affecting skin current much
- Consequences on fine structures during
   TQ are uncertain...

Preliminary Current Quench (CQ) investigations

- Almost all simulations stop converging at some point during the TQ
- However, it is easy and interesting to prolong a TQ simulation in axisymmetric mode
- Here, a large perpendicular heat diffusivity is used
  - Mimics end of TQ
  - Otherwise, localized current sheets (which seem related to local maxima in L<sub>rad</sub>(T))
- Also, need to add impurities in the core, otherwise no radiative collapse (the core even re-heats due to Ohmic heating)

- I<sub>p</sub> decay rate increases with amount of impurities
- When the experimental I<sub>p</sub> decay rate is matched (« 7.5 bar » case), P<sub>rad</sub> is also matched

No surprise because  $P_{rad} = P_{Ohm} \sim d(I_p^2)/dt$  during CQ

Spiky P<sub>rad</sub> in simulations related to treatment of radiation at very low T (sharp cut-off under a certain T, ...)

