

# The role of impurity transport and temperature in MGI induced runaway dynamics

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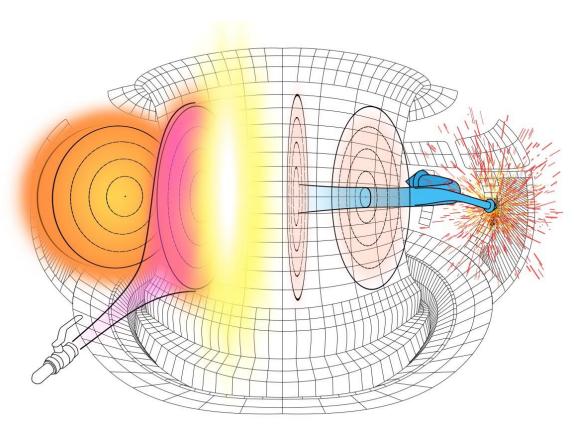
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### Outline

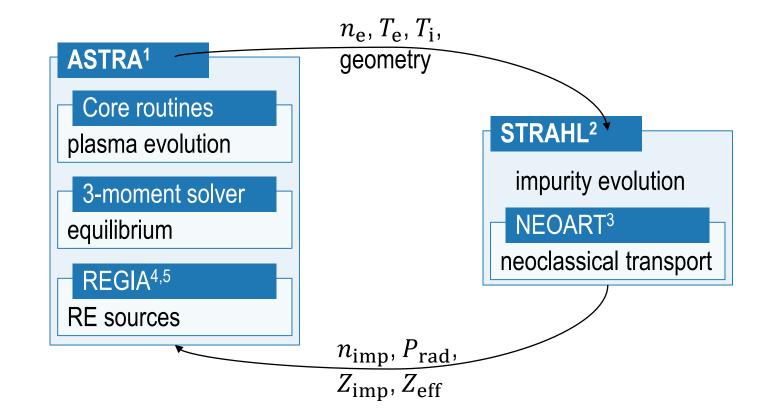
## 1. The transport model ASTRA-STRAHL

- 2. ASDEX Upgrade runaway electron experiments
- 3. Simulating ASDEX Upgrade #33108
  - a. Runaway electron generation
  - b. The role of impurity transport
  - c. Impact of pre-disruption temperature
- 4. Conclusions



## **ASTRA-STRAHL: the coupled transport codes**





<sup>1</sup> Fable et al. *Plasma Phys. Control. Fusion* **55**, <u>074007</u> (2013) <sup>2</sup> Dux et al. *Nucl. Fusion* **39**, <u>1509</u> (1999) <sup>3</sup> Peeters. *Phys. Plasmas* **7**, <u>268</u> (2000) <sup>4</sup> Linder et al. *Nucl. Fusion* **60**, <u>096031</u> (2020) <sup>5</sup> Linder et al. J. Plasma Phys. 87, <u>905870301</u> (2021)

## ASTRA-STRAHL: background plasma evolution with ASTRA

Evolution of plasma quantities Y through macroscopic transport equation<sup>1</sup>

$$\frac{1}{V'}\frac{\partial}{\partial t}(V'Y) = \frac{1}{V'}\frac{\partial}{\partial \rho}\left(V'\langle (\nabla \rho)^2 \rangle \left[D\frac{\partial Y}{\partial \rho} - \nu Y\right]\right) + \sum_i S_i$$

#### Temperatures $T_e$ , $T_i$

- Ohmic heating
- Impurity radiation
- Electron-to-ion heat exchange
- Rapid transport during breakup of magnetic surfaces<sup>2,3</sup> (more on this later)

#### Poloidal flux $\Psi$

- Influenced by RE generation
- $j_{\rm p}$  profile flattened during TQ onset

#### Electron density $n_{\rm e}$

• From quasineutrality



## **ASTRA-STRAHL: impurity evolution with STRAHL**



Evolution of plasma quantities *Y* through macroscopic transport equation

$$\frac{1}{V'}\frac{\partial}{\partial t}(V'Y) = \frac{1}{V'}\frac{\partial}{\partial \rho}\left(V'\langle(\nabla\rho)^2\rangle\left[D\frac{\partial Y}{\partial \rho} - \nu Y\right]\right) + \sum_i S_i$$

#### Impurity densities

- Charge state resolved
- Atomic processes: electron-impact ionization and recombination (coefficients from ADAS<sup>1</sup>)
- Neoclassical transport from NEOART<sup>2</sup>
- Neutral gas propagation at speed of sound
- Rapid transport during breakup of magnetic surfaces<sup>3,4</sup> (more on this later)
- Impurity radiation from line radiation, continuum radiation and ionization losses

## **ASTRA-STRAHL: Runaway Electron Generation In ASTRA (REGIA)**

Evolution of plasma quantities *Y* through macroscopic transport equation

$$\frac{1}{V'}\frac{\partial}{\partial t}(V'Y) = \frac{1}{V'}\frac{\partial}{\partial \rho}\left(V'\langle(\nabla\rho)^2\rangle\left[D\frac{\partial Y}{\partial \rho} - vY\right]\right) + \sum_i S_i$$

#### Runaway electron current density

- Runaway sources S<sub>i</sub> from standalone Fortran module (<u>github.com/o-linder/runawayelectrongeneration</u>)
- Separate populations of RE due to different generation mechanisms
- Average velocity  $\langle v_{\rm RE} \rangle = c$
- No RE losses
- Feed-back on  $\Psi$  evolution

#### Fortran module<sup>1</sup>

#### Dreicer generation

- Classical model by Connor & Hastie<sup>2</sup>
- CODE neural network by Hesslow et al<sup>3</sup>
- Hot-tail generation
  - Model by Smith & Verwichte<sup>4</sup>
- Avalanche generation
  - Classical model by Rosenbluth & Putvinski<sup>5</sup>
  - High-Z model by Hesslow *et al*<sup>6</sup>
- Nuclear generation: not implemented (Recall, application to ASDEX Upgrade)

## **ASTRA-STRAHL: Description of MGI and TQ**



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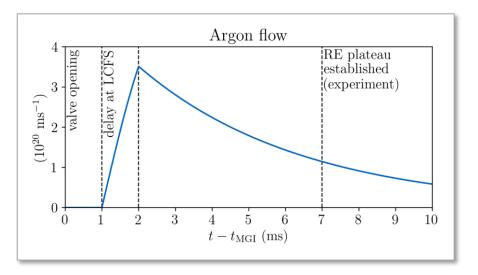
#### Massive gas injection

- Outflow from gas valve described by continuity equation<sup>1</sup>
- Inward propagation with thermal velocity (for Ar):

$$v_{\rm th} = \sqrt{T/m} = 246 \,\mathrm{m/s}$$

- In AUG: valve opens within 1 ms
- In ASTRA: source located 1 cm outside LCFS
   → 1 ms delay

(no need to model propagation from valve to LCFS)

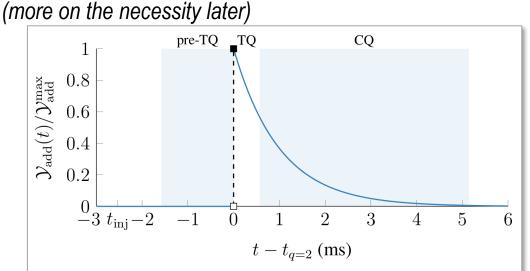


#### Break-up of magnetic surfaces / onset of TQ

- Cold gas front reaches q = 2 surface, triggers (m, n) = (2,1) MHD modes (+ higher harmonics)
- In experiment: flattens  $j_p$  profile, drop in  $l_i$ ,  $I_p$  spike

• In ASTRA<sup>2,3</sup>: 
$$j_p$$
 flattened when  $\left|\frac{\mathrm{d}j_p}{\mathrm{d}\rho}\frac{1}{j_p}\right| > 50$  at  $q = 2$ 

• Additional transport:  $\mathcal{Y}_{add}(t) = \mathcal{Y}_{add}^{\max} \exp\left(-\frac{t-t_{q=2}}{\tau_{add}}\right)$ 





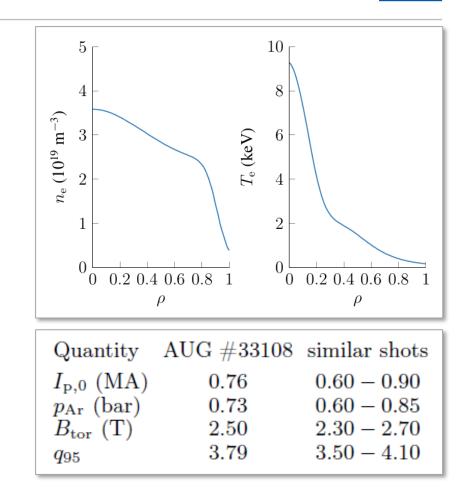
## **ASDEX Upgrade runaway electron experiments**<sup>1-3</sup>

#### MGI in AUG #33108

- Circular L-mode limiter plasma
- Low density  $(3 \times 10^{19} \text{ m}^{-3})$ , high temperature (9 keV)
- Central ECRH (2.6 MW)
- Argon injection (0.73 bar  $\times$  100 cm<sup>3</sup>  $\sim$  7 N<sub>D</sub>)

#### Application

- Used as base case for simulations
- Similar discharges selected for comparison of simulations with experimental trend (impact of  $T_{\rm e}$ )





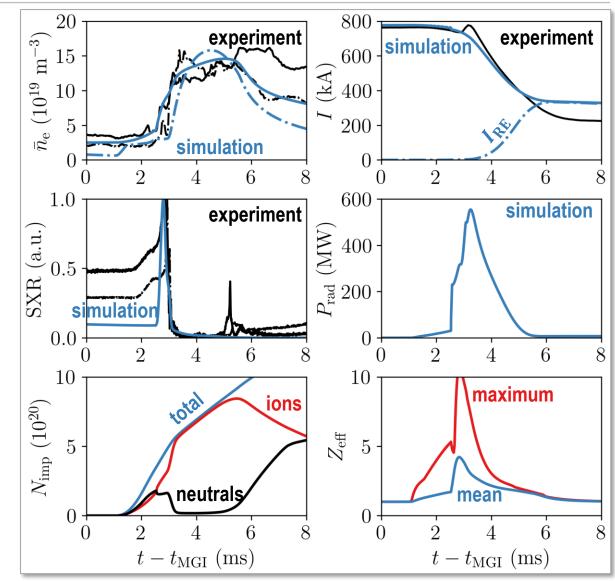
## Simulating ASDEX Upgrade #33108

#### Key experimental observations reproduced

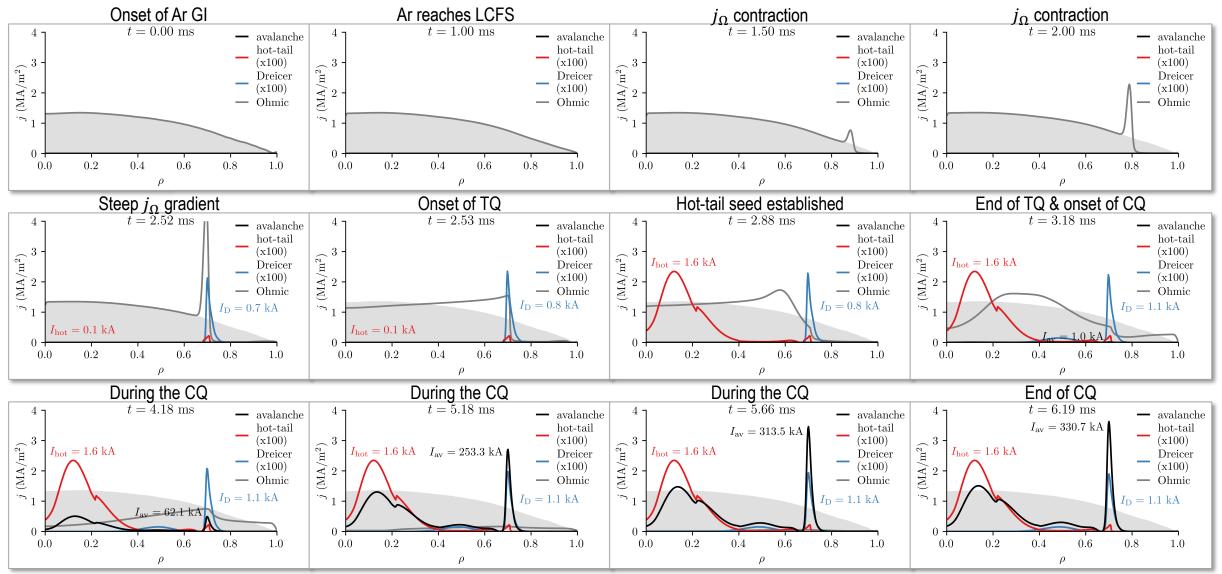
- Increase of electron density  $\bar{n}_{e}$
- Decay of plasma current  $I_{\rm p}$
- Occurrence of TQ

#### Simulation features

- Density increase reproduced
   → current decay reproduced
- Density increase requires additional transport  $D = 100 \text{ m}^2/\text{s}, \quad v = -1000 \text{ m/s},$  $\tau = 1.0 \text{ ms}$
- Thermal energy dissipated by impurity radiation
- Ohmic heating during  $CQ \rightarrow prolonged$  radiation
- Distinct phases of disruption covered (pre-TQ, TQ, CQ)



## **Runaway electron generation: current evolution**



## **Runaway electron generation: contributions**

#### Seed generation

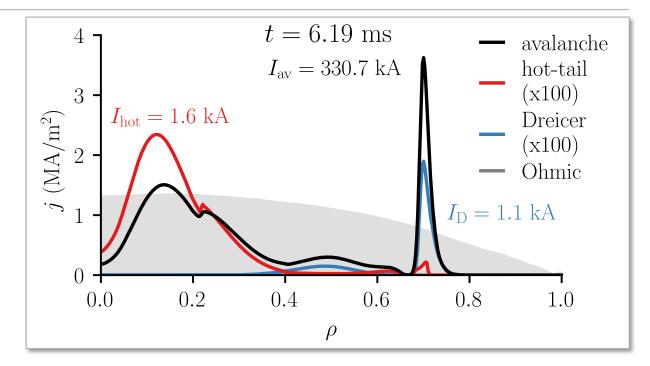
- Only a few kA
- Similar contributions by hot-tail & Dreicer mechanisms

#### Avalanching

- Seed multiplication
- Generates 331 kA of REs
- Final RE beam avalanche dominated
- → RE seed of minor importance (varying strength & composition)

#### **Experimental comparison**

• Higher  $I_{\rm RE}$  simulated due to absence of RE losses







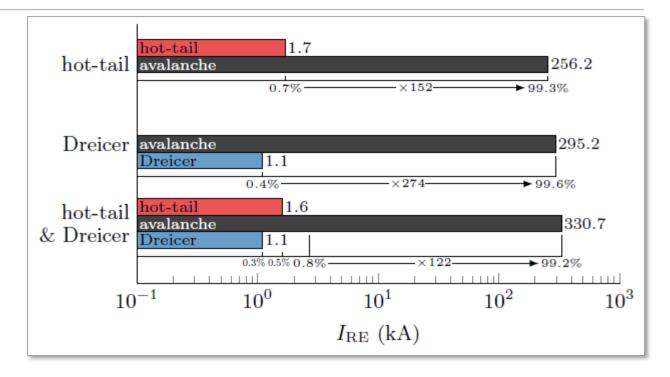
## Runaway electron generation: impact of the RE seed

#### Comparison with only one seed mechanism used

• Reduction of  $I_{RE}^{seed}$  does not affect RE multiplication with equal weight:

hot-tail:  $-37\% I_{RE}^{seed} \rightarrow -23\% I_{av}$ Dreicer:  $-59\% I_{RE}^{seed} \rightarrow -11\% I_{av}$ 

- Decay of  $I_{\Omega}$  at similar time scales, avalanche multiplication time determines post-CQ  $I_{\text{RE}}$  (feedback on  $\Psi$ -evolution)
- → Exact strength of RE seed is of secondary importance due to dominating avalanche generation



## Runaway electron generation: impact of partially ionized impurities

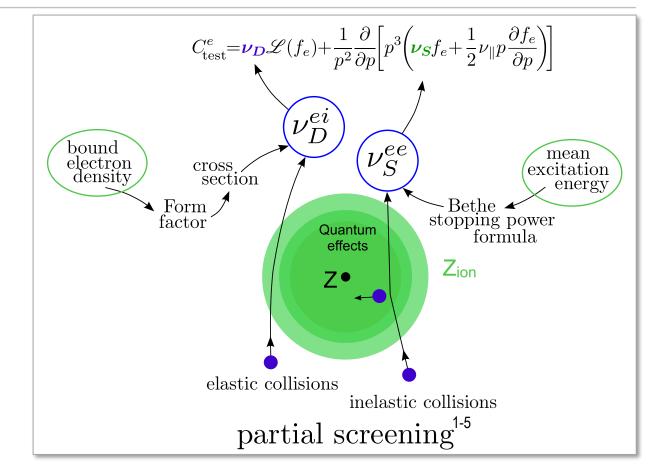
#### Impact of partially ionized impurities<sup>1-5</sup>

- Increased electron-ion friction  $\rightarrow$  hinders runaway
- Relevant in MGI scenarios
- Classical formulae<sup>6-7</sup> assume full ionization
- Effects considered in state-of-the-art models<sup>4-5</sup>

#### ASTRA-STRAHL simulations<sup>8</sup>

• Assess importance of partial screening in selfconsistent simulations:

state-of-the-art<sup>4-5</sup>  $\leftrightarrow$  classical<sup>6-7</sup>



<sup>1</sup> Hesslow et al. Phys. Rev. Lett. 118, <u>255001</u> (2017)
 <sup>2</sup> Hesslow et al. Plasma Phys. Control. Fusion 60, <u>074010</u> (2018)
 <sup>3</sup> Hesslow et al. J. Plasma Phys. 84, <u>905840605</u> (2018)

<sup>4</sup> Hesslow et al. Nucl. Fusion **59**, <u>084004</u> (2019)
 <sup>5</sup> Hesslow et al. J. Plasma Phys. **85**, <u>475850601</u> (2019)
 <sup>6</sup> Connor et al. Nucl. Fusion **15**, <u>415</u> (1975)

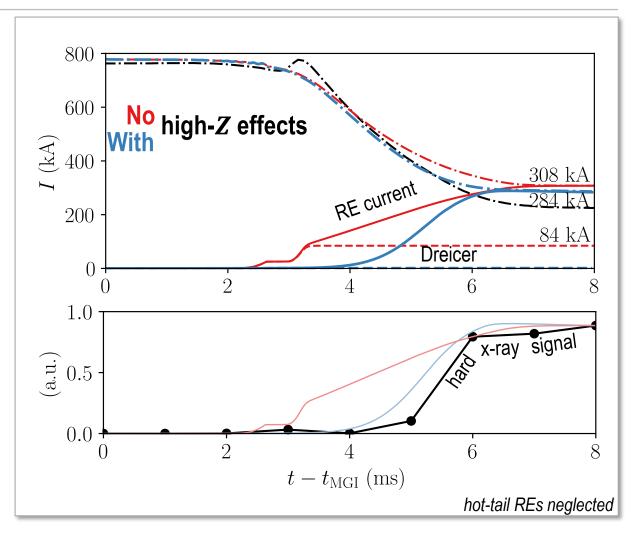
<sup>*<u>7</u></sup> Rosenbluth <i>et al. Nucl. Fusion* **37**, <u>1355</u> (1997) <sup><u>8</u></sup> Linder *et al. Nucl. Fusion* **60**, <u>096031</u> (2020)</sup>



## **Runaway electron generation: model validation**

#### Absence of high-Z effects

- Dreicer generation overestimated (earlier onset & stronger, 84 kA)
- Avalanche multiplication reduced (slower rise of RE current during CQ)
- Decay of total current slowed down
   → less Ohmic heating, less radiation
- Final  $I_{\rm RE}$  similar, but different composition
- Hard x-ray signal: High  $I_{\rm RE}$  only at end of CQ
- → Simulations consistent with experiment only when considering high-*Z* effects
- $\rightarrow$  High-Z interactions important for runaway!



## The role of impurity transport: impact of transport mechanisms

#### **Mechanisms considered**

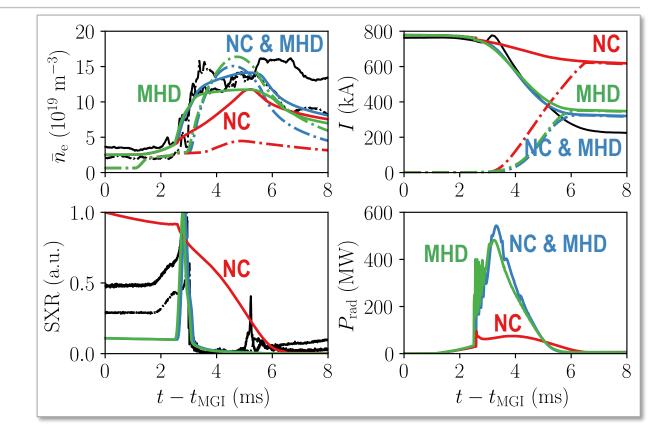
- Rapid redistribution (MHD effects due to breakup of magnetic surfaces)
- Neoclassical effects

#### Absence of rapid redistribution

- Impurity propagation driven by neutral gas
- Increase of electron density  $\bar{n}_{e}$  not matched
- $\rightarrow$  Much slower TQ!

#### Absence of neoclassical effects

- Inward transport less effective
- → Rapid redistribution & neoclassical effects relevant for impurity transport



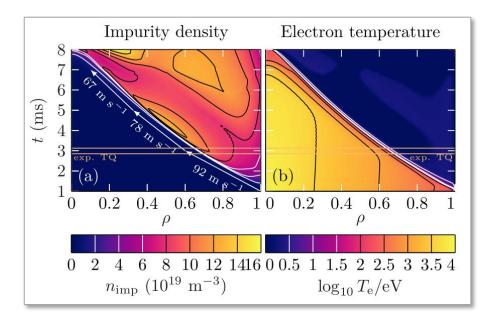


## The role of impurity transport: rapid redistribution



#### Absence of rapid redistribution

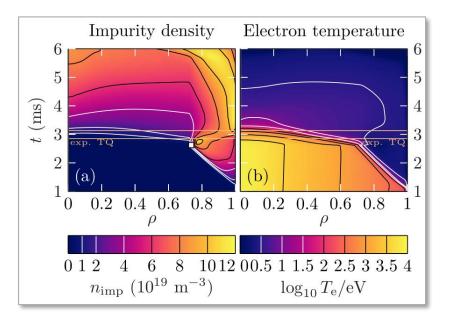
- Impurity propagation driven by neutral gas
- Slow TQ over several ms



#### Considering rapid redistribution

- Central impurity density increases during CQ
- TQ on experimental sub-ms time scales
- Note, only order of magnitude values used:  $D = 100 \text{ m}^2/\text{s}, \quad v = -1000 \text{ m/s},$  $\tau = 1.0 \text{ ms}$

(variation by around 50% describes experiment adequately)



## The role of impurity transport: neoclassical transport

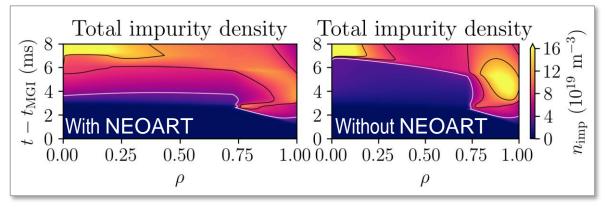
#### In absence of additional transport

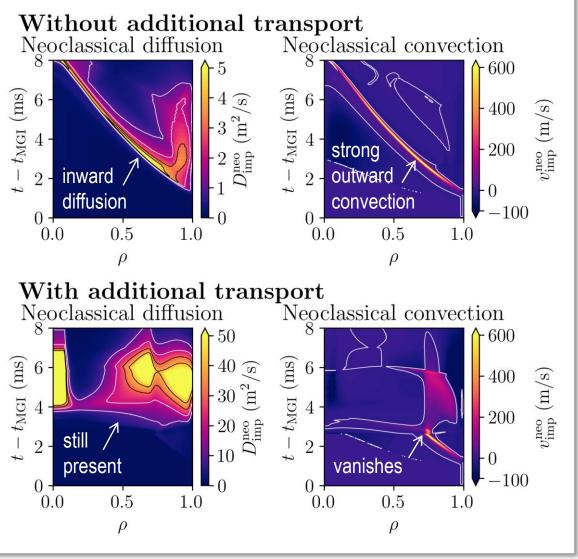
- Diffusion & strong outward convection almost cancel
- Propagation driven by neutrals
- Slow inward propagation of material

#### With additional transport

- Outward convection vanishes; diffusion present
- Neoclassical transport contributes noticeably to inward transport
   (ourrent decay too slow in absonce of neoclassical off

(current decay too slow in absence of neoclassical effects)







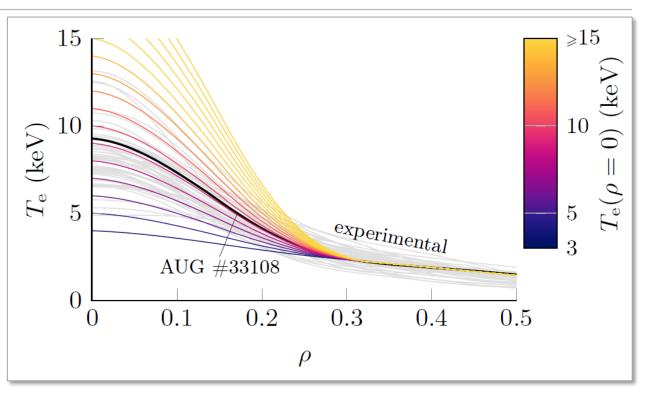
## Impact of pre-disruption temperature: simulation setup

## Setup

- In AUG experiments, on-axis ECRH during last 0.1 ms prior to MGI
- Scale ECRH contribution of  $T_e(\rho)$  in AUG #33108

$$T_{\rm e}(\rho) = \frac{T_{\rm tar} - T_{\rm bg}(0)}{T_{\rm ECRH}(0)} T_{\rm ECRH}(\rho) + T_{\rm bg}(\rho)$$
  
target background

- Temperature unaffected for ho > 0.35



## Impact of pre-disruption temperature: the RE seed



#### **Dreicer population**

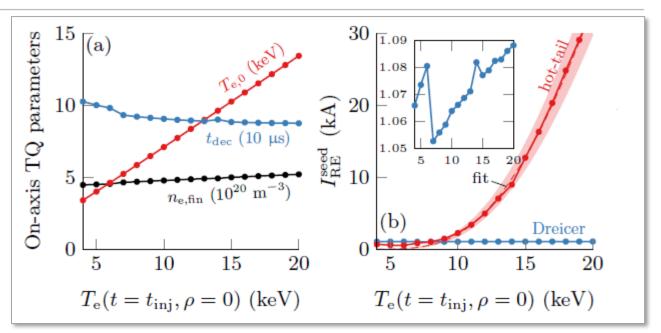
- Virtually unaffected:  $I_{\rm D} = 1.1 \text{ kA}$
- No (significant)  $T_{e}$  change in region of strongest generation

#### Hot-tail population

- Grows exponentially  $I_{hot}$ : 0.6 kA  $\rightarrow$  33.7 kA
- Strongest increase above 10 keV
- Described by exponential

$$I_{\rm hot}^{\rm fit}(T_{\rm e,0}(\rho=0)) = (914 \pm 58) \exp\left(-4\left\{\frac{\tilde{\nu}\ln\Lambda(t_0)\,\langle n_{\rm e,fin}t_{\rm dec}\rangle}{T_{\rm e,0}(\rho=0)^{3/2}}\right\}^{2/3}\right) \,\,\mathrm{kA}$$

- Strong increase of temperature (by design), but density not at similar rate
- $\rightarrow$  Strong increase of hot-tail runaway for hotter plasmas



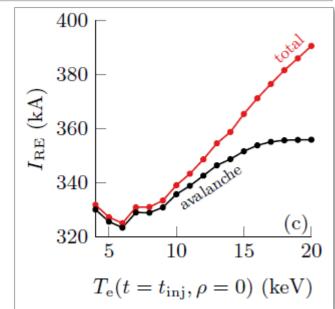
## Impact of pre-disruption temperature: avalanche & total current

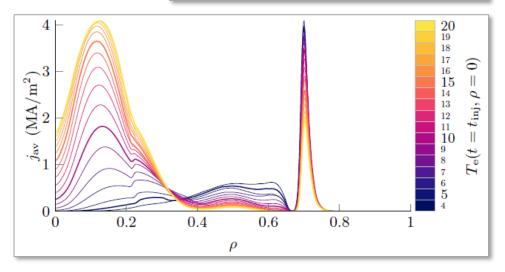
#### **Multiplication**

- Avalanche (& total) RE current ~ 330 kA for  $T_{\rm e}$  < 9 keV
- Between 9 keV and 17 keV,  $I_{av}$  increases (additional multiplication between 2 3)
- Avalanche current saturates for  $T_{\rm e} > 17~{\rm keV}$  at 356 kA
- Radial distribution of  $j_{\rm av}$  changes (as hot-tail becomes more important): mid radius &  $q=2 \rightarrow \rho \sim 0.12$

#### **Total current**

- For  $T_{\rm e} > 9 \, {\rm keV}$ , grows linearly due to hot-tail contribution
- At  $T_{\rm e} = 20 \text{ keV}$ , hot-tail constitutes 9%  $I_{\rm RE}$
- → Relative importance of multiplication decreases due to finite poloidal magnetic flux available (less avalanching in ITER?)







## Impact of pre-disruption temperature: experimental comparison

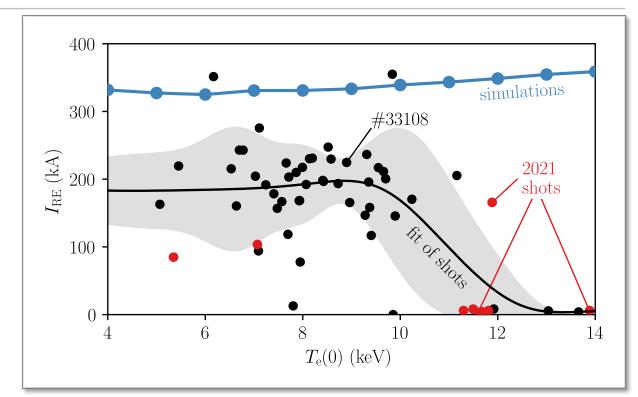


#### For $T_{\rm e} < 9 \; {\rm keV}$

- RE current ~ constant (experiment & simulation)
- Simulated I<sub>RE</sub> larger due to absence of RE losses

For  $T_{\rm e} > 9 \; {\rm keV}$ 

- In simulation, *I*<sub>RE</sub> increasing (hot-tails)
- None in experiment at  $T_e > 12 \ keV$
- In 2021 campaign, one discharge with RE beam close to 12 keV
- Reason: Loss of RE seed during breakup of magnetic surfaces?



## **Further reading**



#### More details on self-consistent ASTRA-STRAHL simulations of Ar MGI in ASDEX Upgrade #33108 are presented in the following publications:

**Physics** 

Abstract

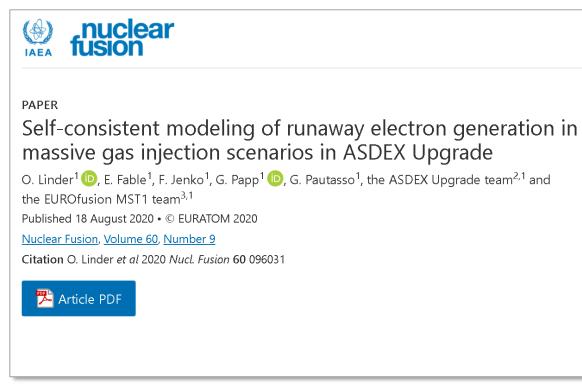
Introduction

Model description

ASDEX Upgrade

evnerimente

runaway electron



Linder et al. Nucl. Fusion 60, 096031 (2020) https://doi.org/10.1088/1741-4326/ab9dcf



The formation of a substantial postdisruption runaway electron current in ASDEX Upgrade material injection experiments is determined by avalanche multiplication of a small seed population of runaway electrons. For the investigation of these scenarios, the runaway electron description of the coupled 1.5-D transport solvers ASTRA-STRAHL is amended by a fluid model describing electron runaway caused by the hot-tail mechanism. Applied in

#### Linder et al. J. Plasma Phys. 87, 905870301 (2021) https://doi.org/10.1017/S0022377821000416

## Conclusions



- 1. Successful ASTRA-STRAHL simulations of RE dynamics on ASDEX Upgrade (background plasma, MGI, RE generation)
- 2. High-*Z* effects important for RE generation
- 3. Impurity transport due to MHD & neoclassical effects
- 4. Discrepancies w.r.t. experiment in high  $T_{e}$  scenarios suggest seed RE loss



- CONNOR, J.W. and HASTIE, R.J.. Relativistic limitations on runaway electrons. *Nucl. Fusion* **15**, <u>415</u> (1975)
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- FABLE, E., ANGIONI, C., IVANOV, A.A. et al. Dynamical coupling between magnetic equilibrium and transport in tokamak scenario modelling, with application to current ramps. *Plasma Phys. Control. Fusion* **55**, <u>074007</u> (2013)
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- HESSLOW, L., EMBREUS, O., WILKIE, G.J. et al. Effect of partially ionized impurities and radiation on the effective critical electric field for runaway generation. *Plasma Phys. Control. Fusion* **60**, <u>074010</u> (2018)
- HESSLOW, L., EMBREUS, O., HOPPE, M. et al. Generalized collision operator for fast electrons interacting with partially ionized impurities. *J. Plasma Phys.* **84**, <u>905840605</u> (2018)
- HESSLOW, L., EMBREUS, O., VALLHAGEN, O. et al. Influence of massive material injection on avalanche runaway generation during tokamak disruptions. *Nucl. Fusion* **59**, <u>084004</u> (2019)
- HESSLOW, L., UNNERFELT, L., VALLHAGEN, O. et al. Evaluation of the Dreicer runaway growth rate in the presence of high-*Z* impurities using a neural network. *J. Plasma Phys.* **85**, <u>475850601</u> (2019)

- LINDER, O., FABLE, E., JENKO, F. et al. Self-consistent modeling of runaway electron generation in massive gas injection scenarios in ASDEX Upgrade. *Nucl. Fusion* **60**, <u>096031</u> (2020)
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- SMITH, H.M. & VERWICHTE, E.. Hot tail runaway electron generation in tokamak disruptions. *Phys. Plasmas* **15**, <u>072502</u> (2008)

SUMMERS, H.P. The ADAS User Manual, version 2.6. <u>https://www.adas.ac.uk</u>.



## **Appendix: Average runaway electron velocity**

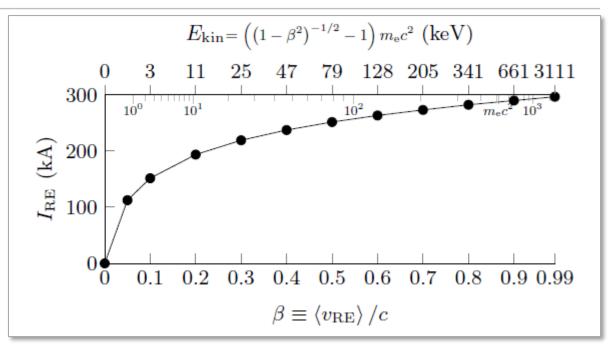
#### Impact of assumption $\langle v_{RE} angle = c$

- Valid for  $E_{kin} \ge 3.1 \text{ MeV}$ ( $\le 1\%$  deviation)
- Conditions not fulfilled during early stages
- Yet, reduction of  $\langle v_{RE} \rangle$  to below c  $\rightarrow$  slight decrease of  $I_{RE}$ (only assuming  $E_{kin} \sim T_{e,0}$ : strong decrease)
- Reduction of  $\langle v_{RE} \rangle \equiv$  reduction of seed population! *(mathematically)*

 $\frac{\partial j_{\text{seed}}}{\partial t} = e \left\langle v_{\text{RE}} \right\rangle S_{\text{seed}} = e \left( c\beta \right) S_{\text{seed}} = ec \left( \beta S_{\text{seed}} \right)$ 

$$\frac{\partial j_{\rm av}}{\partial t} = e \left\langle v_{\rm RE} \right\rangle n_{\rm RE} \tilde{S}_{\rm av} = \left( j_{\rm av} + j_{\rm seed} \right) \tilde{S}_{\rm av}$$

→ Avalanche contribution not affected proportionally; total current only somewhat reduced (same argument as before!)





## Appendix: Off-axis hot-tail density peaking

#### Important parameters for hot-tail generation<sup>1</sup>

Simple estimate<sup>2</sup>

1

$$n_{\text{hot}}^{\text{simple}}(t_{\text{fin}}) \propto n_{\text{e},0} \exp\left(-4 \,\tilde{\nu} \frac{n_{\text{e},\text{fin}}^{2/3} t_{\text{dec}}^{2/3}}{T_{\text{e},0}}\right)$$

- Pre-disruption temperature  $T_{e,0}$
- Decay time scale  $t_{dec}$
- Post-disruption density  $n_{e,fin}$
- → Off-axis hot-tail peaking due to on-axis density peaking post-TQ

